

BMS₂

Black holes — BMS & Integrability

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with Afshar, González, Vassilevich 1911.05739

Outline

Motivation

Kinematics

Dynamics

Relation to SYK/JT

Outlook

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Asymptotic symmetries

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- ▶ basic ingredient of AdS/CFT tests based on symmetries
- ▶ captures universal UV features of QFTs (conformal symmetries)
- ▶ Brown–Henneaux precursor for AdS₃/CFT₂

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▶ Holography beyond AdS/CFT

- ▶ asymptotic holography beyond AdS/CFT?
- ▶ near horizon holography?
- ▶ asymptotic symmetries important input for structure of dual QFT

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- ▶ (extended) BMS₄ algebra ($J_a(x)$: diff S^2 or restriction thereof)

$$\{J_a(x), J_b(x')\} = (J_a(x')\partial_b - J_b(x)\partial'_a) \delta(x - x')$$

$$\{J_a(x), P(x')\} = \left(\frac{s}{2} P(x')\partial_a - P(x)\partial'_a\right) \delta(x - x')$$

$$\{P(x), P(x')\} = 0$$

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- ▶ get same algebra as near horizon symmetries (in any dimension ≥ 3)
Donnay, Giribet, González, Pino '15 $s = 0$ ('scalar super-translations')
DG, Perez, Troncoso, Sheikh-Jabbari, Zwickel '19 arbitrary s

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- ▶ Contraction means $\ell \rightarrow \infty$ and yields BMS₃ (M_n : super-translations)
- ▶ Example: Einstein gravity

$$c = \bar{c} = \frac{3\ell}{2G} \quad \Rightarrow \quad c_L = 0 \quad c_M = \frac{3}{G}$$

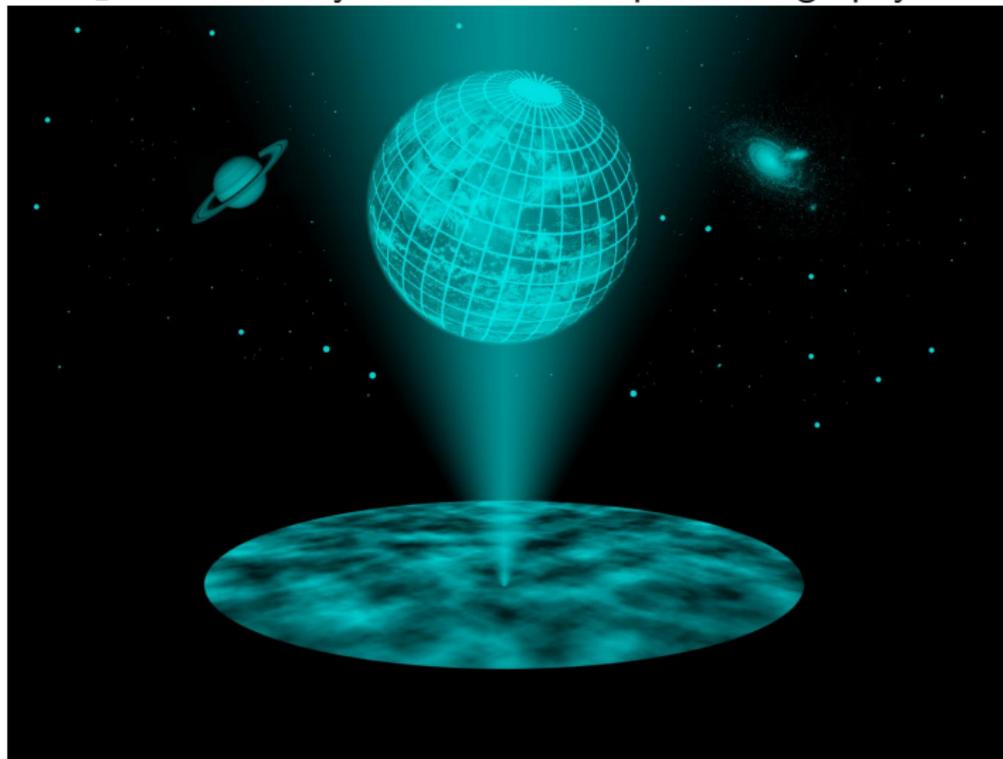
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Motivation for BMS_2

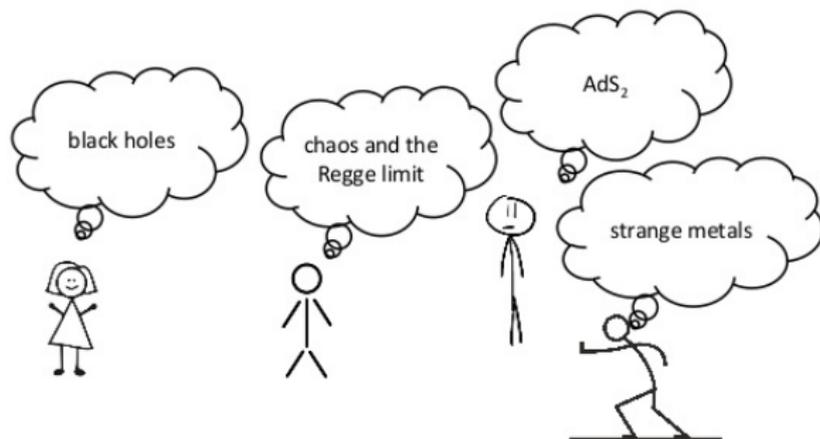
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- ▶ because it is there (maybe)
- ▶ BMS_2 useful for toy models of flat space holography
- ▶ BMS_2 perhaps useful for near horizon holography
- ▶ construct SYK-like models with asymptotically flat gravity side

The SYK model is a **strongly interacting** quantum system that is **solvable** at large N .



slide from Stanford's talk at Strings 2017

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Ignore difficulties and proceed*

* van Nieuwenhuizen: task of theoretical physicists is to break no-go theorems

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Asymptotically Ricci-flat metrics

- ▶ Gauge-fix to Eddington–Finkelstein coordinates

$$ds^2 = -2 du dr + K(u, r) du^2$$

Not obvious that this is possible with proper gauge trafos!
Same remark applies to *any* gauge fixing, e.g. in AdS₃

Asymptotically Ricci-flat metrics

- ▶ Gauge-fix to Eddington–Finkelstein coordinates

$$ds^2 = -2 du dr + K(u, r) du^2$$

- ▶ Demand Ricci-flatness

$$K(u, r) = 2\mathcal{P}(u) r + 2\mathcal{T}(u)$$

Note: for constant \mathcal{P} and \mathcal{T} Killing horizon

$$r_h = -\frac{\mathcal{T}}{\mathcal{P}}$$

Assume in most of talk constant \mathcal{P} and \mathcal{T}

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- ▶ Whatever the gravity theory is going to be, require the following boundary conditions for metric

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Determine next asymptotic Killing vectors

Asymptotic Killing vectors

- ▶ Class of metrics

$$ds^2 = -2 du dr + 2 (\mathcal{P}(u) r + \mathcal{T}(u)) du^2$$

preserved by asymptotic Killing vectors

$$\xi(\epsilon, \eta) = \epsilon(u)\partial_u - (\epsilon'(u)r + \eta(u))\partial_r$$

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- ▶ Metric functions transform non-trivially

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- ▶ Looks promising!

\mathcal{P} like $u(1)$ current

\mathcal{T} like Virasoro generator

- ▶ Lie-bracket algebra of asymptotic Killing vectors

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Call this algebra BMS₂

Can (and will) have non-trivial central extensions

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Dismiss winding mode and focus on warped Witt algebra

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$$I = \frac{1}{16\pi G_2} \int_{\mathcal{M}} d^2x \sqrt{|g|} [XR - U(X)(\nabla X)^2 - V(X)] \\ - \frac{1}{8\pi G_2} \int_{\partial\mathcal{M}} dx \sqrt{|\gamma|} [XK - S(X)] + I^{(m)}$$

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- ▶ Interesting option: couple 2d dilaton gravity to **matter**

Selected list of models (see review [hep-th/0604049](https://arxiv.org/abs/hep-th/0604049))

Black holes in (A)dS₂, asymptotically flat or arbitrary spaces (Wheeler property)

Model	$U(X)$	$V(X)$
1. Schwarzschild (1916)	$-\frac{1}{2X}$	$-\lambda^2$
2. Jackiw-Teitelboim (1984)	0	ΛX
3. Witten Black Hole (1991)	$-\frac{1}{X}$	$-2b^2 X$
4. CGHS (1992)	0	-2Λ
5. (A)dS ₂ ground state (1994)	$-\frac{a}{X}$	BX
6. Rindler ground state (1996)	$-\frac{a}{X}$	BX^a
7. Black Hole attractor (2003)	0	BX^{-1}
8. Spherically reduced gravity ($N > 3$)	$-\frac{N-3}{(N-2)X}$	$-\lambda^2 X^{(N-4)/(N-2)}$
9. All above: ab -family (1997)	$-\frac{a}{X}$	BX^{a+b}
10. Liouville gravity	a	$be^{\alpha X}$
11. Reissner-Nordström (1916)	$-\frac{1}{2X}$	$-\lambda^2 + \frac{Q^2}{X}$
12. Schwarzschild-(A)dS	$-\frac{1}{2X}$	$-\lambda^2 - \ell X$
13. Katanaev-Volovich (1986)	α	$\beta X^2 - \Lambda$
14. BTZ/Achucarro-Ortiz (1993)	0	$\frac{Q^2}{X} - \frac{J}{4X^3} - \Lambda X$
15. KK reduced CS (2003)	0	$\frac{1}{2} X(c - X^2)$
16. KK red. conf. flat (2006)	$-\frac{1}{2} \tanh(X/2)$	$A \sinh X$
17. 2D type 0A string Black Hole	$-\frac{1}{X}$	$-2b^2 X + \frac{b^2 q^2}{8\pi}$
18. exact string Black Hole (2005)	lengthy	lengthy

Gauge theoretic formulation as Poisson-sigma model (PSM)

- ▶ 2d analogue of Chern–Simons formulation of 3d gravity: PSM
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- ▶ still need to choose gauge algebra and bilinear form

Cangemi–Jackiw version of Callan–Giddings–Harvey–Strominger

- ▶ Choose Maxwell algebra

$$[P_a, P_b] = \epsilon_{ab} Z \qquad [P_a, J] = \epsilon_a{}^b P_b$$

with bilinear form

$$\langle J, Z \rangle = -1 \qquad \langle P_a, P_b \rangle = \eta_{ab}$$

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$$R = 0 \quad \Rightarrow \quad \text{Ricci-flat}$$

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- ▶ translate our bc's into BF-formulation

Boundary conditions in BF formulation

- ▶ Ansatz (worked nicely for Jackiw–Teitelboim; inspired by 3d)

$$\mathcal{A} = b^{-1}(\mathrm{d}+a)b \qquad B = b^{-1}xb$$

with group element $b = \exp(-r P_+)$ and

$$a = (\mathcal{T}(u)P_+ + P_- + \mathcal{P}(u)J) \mathrm{d}u$$
$$x = x^+(u)P_+ + x_1(u)P_- + YJ + x_0(u)Z$$

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- ▶ yields metric shown before, dilaton

$$X = x_1(u)r + x_0(u)$$

and Maxwell field $A = r \mathrm{d}u$

get BMS_2 asymptotic symmetries!

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and Maxwell field $A = r du$

- ▶ Maxwell field preserved by combined diffeos and gauge trafos

$$\delta A_\nu = \xi^\mu \partial_\mu A_\nu + A_\mu \partial_\nu \xi^\mu + \partial_\nu \sigma \quad \xi(\epsilon, \eta) = \epsilon(u)\partial_u - (\epsilon'(u)r + \eta(u))\partial_r$$

provided $\eta = \sigma'$

either η has no 0-mode or σ not single-valued (winding modes)

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- ▶ focus on case $\delta_\sigma \oint A = 0$ (no winding modes) \Rightarrow warped Witt algebra

Twisted warped boundary action (see also Afshar '19)

- ▶ Variation of Euclidean BF action ($t = iu$)

$$\delta I_{\text{BF}} = \text{bulk-EOM} - \kappa \oint dt \langle x, \delta a_t \rangle$$

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- ▶ defining $1/x_1 \sim \partial_t f$ and $x_0/x_1 \sim \partial_t g$ result is

$$I_{\text{tw}}[h, g] = \kappa \int_0^\beta d\tau \left(\mathcal{T} h'^2 - g' \left(i\mathcal{P} h' + \frac{h''}{h'} \right) \right)$$

with $\tau := f(t)$, $h(\tau) := -f^{-1}(\tau)$ and $\tau \sim \tau + \beta$ (prime means $d/d\tau$)

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twisted warped action is flat space analogue of Schwarzian action!

- ▶ Schwarzian action: group action for Virasoro coadjoint orbits
- ▶ twisted warped action: group action for twisted warped coadjoint orbits

$$[L_n, L_m] = (n - m) L_{n+m}$$

$$[L_n, J_m] = -m J_{n+m} - i\kappa(n^2 - n) \delta_{n+m, 0}$$

$$[J_n, J_m] = 0$$

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Twisted warped action resembles effective action for complex SYK

Outline

Motivation

Kinematics

Dynamics

Relation to SYK/JT

Outlook

- ▶ twisted warped Hamiltonian action

$$I_{\text{tw}} = -\kappa \int_0^\beta dt (p_i \dot{q}_i - p_1 p_2 - e^{q_1} p_3) \quad i = 1, 2, 3$$

where $q_3 = \exp(i\mathcal{P}h)$ and $q_2 = g + ih\mathcal{T}/\mathcal{P}$ (rest: auxiliary fields)

Hamiltonian formulation

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- ▶ solutions

$$q_3 = h_0 + h_1 e^{i\tau/\tau_0} \quad q_2 = g_0 - ig_1 \tau + g_2 e^{i\tau/\tau_0}$$

five integration constants g_0, g_1, g_2, h_0, h_1 ; periodicity $\tau_0 = \beta/(2\pi)$

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- ▶ on-shell action $I_{\text{tw}}|_{\text{EOM}} = -2\pi\kappa g_1$

- ▶ Assuming g_1 independent from temperature get entropy

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- ▶ useful property for scaling limit from complex SYK

- ▶ Effective action governing collective low T modes of complex SYK

$$I_{\text{cSYK}}[h, g] = \frac{NK}{2} \int_0^\beta d\tau \left(g' + \frac{2\pi i \mathcal{E}}{\beta} h' \right)^2 - \frac{N\gamma}{4\pi^2} \int_0^\beta d\tau \left\{ \tan \left(\frac{\pi}{\beta} h \right); \tau \right\}$$

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$\{f; \tau\} := f'''/f' - \frac{3}{2}(f''/f')^2$ Schwarzian derivative

N (large) number of complex fermions

$N\gamma$ specific heat at fixed charge

K zero temperature compressibility

\mathcal{E} spectral asymmetry parameter

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- ▶ according to our thermodynamics need limit $N\gamma \rightarrow \infty$
(infinite specific heat)

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- ▶ inserting these limits into $I_{\text{cSYK}}[h, g]$ yields twisted warped action

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with $\kappa^2 \sim N^a \gamma K$ kept finite

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- ▶ then take limit $\hat{K} \rightarrow 0$, $c \rightarrow \infty$ keeping fixed $\kappa = \sqrt{-\frac{c\hat{K}}{24}}$

Conclusions

For more details see [Afshar, González, DG, Vassilevich 1911.05739](#)

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Outline

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Kinematics

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Relation to SYK/JT

Outlook

Future developments

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Thanks for your attention!

