

Stability of the Boltzmann equation

Take renormalization $\beta_\delta(z) = \frac{z}{1+\delta z}$ (concave), $\beta'_\delta(z) = \frac{1}{(1+\delta z)^2}$ (convex)

$$\left. \begin{aligned} \beta_\delta(f^n) &\xrightarrow{*} \beta_\delta \leq \beta_\delta(f) \\ \beta'_\delta(f^n) &\xrightarrow{*} h_\delta \geq \beta'_\delta(f) \\ \beta'_\delta(f^n) f^n = \frac{f^n}{(1+\delta f^n)^2} = \beta_\delta(f^n) (1 - \delta \beta_\delta(f^n)) &\xrightarrow{*} g_\delta \leq \beta_\delta (1 - \delta \beta_\delta) \end{aligned} \right\} \begin{array}{l} \text{in weak}^* \\ \text{top. of} \\ L_{loc}^\infty(dt dx dv) \end{array}$$

$$(\partial_t + v \cdot \nabla_x) (\beta_\delta(f^n)) = \beta'_\delta(f^n) C(f^n) \xrightarrow{\mathcal{Q}'} (\partial_t + v \cdot \nabla_x) \beta_\delta = h_\delta C^+(f) - g_\delta f + \left[\int b(\cdot, \sigma) d\sigma \right]$$

One can show that $\beta_\delta, g_\delta \rightarrow f$ as $\delta \rightarrow 0$ in $L_{loc}^1(dt dx dv)$
and $h_\delta \rightarrow 1$ as $\delta \rightarrow 0$ in $L_{loc}^p(dt dx dv)$, $\forall 1 \leq p < \infty$.

Take an admissible renormalization $\beta(z)$

$$\Rightarrow (\partial_t + v \cdot \nabla_x) \beta(\beta_\delta) = \beta'(\beta_\delta) h_\delta C^+(f) - \beta'(\beta_\delta) g_\delta f + \left[\int b(\cdot, \sigma) d\sigma \right]$$

$$\text{Let } \delta \rightarrow 0 \Rightarrow (\partial_t + v \cdot \nabla_x) \beta(f) = \beta'(f) C^+(f) - \beta'(f) f + \left[\int b(\cdot, \sigma) d\sigma \right]$$

Deriving Navier-Stokes Equations from the Boltzmann Equation

Nondimensional form of the Boltzmann equation

$$St \partial_t f + v \cdot \nabla_x f = \frac{1}{Kn} C(f)$$

Kn = Knudsen number $\sim$ mean free path

St = Strouhal number $\sim$ time scaling

Hydrodynamic regimes happen when $Kn \rightarrow 0$. Think of it as a continuum limit.

Further consider a global normalized equilibrium state

$$M(v) = \frac{1}{(2\pi)^{d/2}} e^{-\frac{v^2}{2}}$$

and its fluctuations $f = M(1 + Ma g)$.

Ma = Mach number $\sim$ ratio of bulk velocity to speed of sound.

Compressible Euler regime

$$(\partial_t + v \cdot \nabla_x) f_\varepsilon = \frac{1}{\varepsilon} C(f_\varepsilon).$$

$$\varepsilon \rightarrow 0 \Rightarrow f_\varepsilon \rightarrow f \text{ and } C(f) = 0 \Rightarrow f = \frac{\rho}{(2\pi\theta)^{d/2}} e^{-\frac{|v-u|^2}{2\theta}}$$

Conservation laws of mass, momentum and energy become the compressible Euler system in the limit:

$$\begin{cases} \partial_t \rho + \operatorname{div}(\rho u) = 0 \\ \partial_t(\rho u) + \operatorname{div}(\rho u \otimes u) + \nabla(P\theta) = 0 \\ \partial_t\left(\frac{1}{2}\rho u^2 + \frac{d}{2}P\theta\right) + \operatorname{div}\left(\left(\frac{1}{2}\rho u^2 + \frac{d+2}{2}P\theta\right)u\right) = 0 \end{cases}$$

Note: von Kármán relation (physics principle) $Kn \sim \frac{Ma}{Re}$ (Re is the Reynolds number) tells us that $Re \rightarrow \infty$ as $Kn \rightarrow 0$ if $Ma \sim 1$.
 $\Rightarrow$ compressible Navier-Stokes cannot be derived with such hydrodynamic limits.

Acoustic waves regime

$$(\partial_t + v \cdot \nabla_x) f_\varepsilon = \frac{1}{\varepsilon} C(f_\varepsilon), \quad f_\varepsilon = \Pi(1 + \varepsilon^m g_\varepsilon), \quad m > 0.$$

$$g_\varepsilon \rightarrow g = p + v \cdot u + \left(\frac{v^2}{2} - \frac{d}{2}\right) \theta \quad (\in \text{kernel of linearized Boltzmann operator})$$

$$\begin{cases} \partial_t p + \operatorname{div} u = 0 \\ \partial_t u + \nabla(p + \theta) = 0 \\ \frac{d}{2} \partial_t(p + \theta) + \frac{d+2}{2} \operatorname{div} u = 0 \Rightarrow \frac{d}{2} \partial_t \theta + \operatorname{div} u = 0 \end{cases}$$

Incompressible regimes (Bardos, Golse, Levermore, 1991)

$$(\varepsilon \partial_t + v \cdot \nabla_x) f_\varepsilon = \frac{1}{\varepsilon^q} C(f_\varepsilon), \quad f_\varepsilon = \Pi(1 + \varepsilon^m g_\varepsilon)$$

- $q=1, m=1$, incompressible Navier-Stokes-Fourier system (details later)
- $q>1, m=1$, incompressible Euler-Fourier system
- $q=1, m>1$, incompressible Stokes-Fourier system.

equation for fluctuations:

$$(\varepsilon \partial_t + v \cdot \nabla_x) g_\varepsilon = \underbrace{\frac{1}{\varepsilon^q} \int (g'_\varepsilon + g'_{\varepsilon*} - g_\varepsilon - g_{\varepsilon*}) \Pi_* b dv_* d\sigma}_{\frac{1}{\varepsilon^{q+m}} \Pi C(f_\varepsilon)} + \underbrace{\frac{1}{\varepsilon^{q-m}} \int (g'_\varepsilon g'_{\varepsilon*} - g_\varepsilon g_{\varepsilon*}) \Pi_* b dv_* d\sigma}_{Q(g_\varepsilon)}$$

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Suppose $g_\varepsilon \rightarrow g$.

infinitesimal Maxwellian

Order $\frac{1}{\varepsilon}$: $L(g)=0 \Rightarrow g' + g'_* = g + g_* \Rightarrow \boxed{g = \rho + v \cdot u + \left(\frac{v^2}{2} - \frac{d}{2}\right)\theta}$

$\varepsilon \rightarrow 0$ in moment equations: $\int (v \cdot \nabla_x g) \begin{pmatrix} 1 \\ v \\ v^2 \end{pmatrix} M dv = 0$

$\Rightarrow \boxed{\operatorname{div} u = 0} \quad \text{and} \quad \boxed{\nabla(\rho + \theta) = 0}$

- Evolution equations. Difficulty: fluxes are singular.

Note: $L: L^2(M dv) \rightarrow L^2(M dv)$ (if $b \equiv 1$, otherwise need to use weights in velocity)

$L = \operatorname{Id} - (\text{compact})$ if $b \equiv 1$

is self-adjoint and Fredholm of index zero on $L^2(M dv)$.

$\Rightarrow \operatorname{Ran} L$ is exactly the orthogonal complement of $\operatorname{Ker} L$.

Moreover $\operatorname{Ker} L = \operatorname{span}\{1, v_1, \dots, v_d, |v|^2\}$.

Introduce the hydrodynamic projection Πg_ε of g_ε onto $\operatorname{Ker} L$:

$$\Pi g_\varepsilon = \rho_\varepsilon + u_\varepsilon \cdot v + \theta_\varepsilon \left(\frac{v^2}{2} - \frac{d}{2} \right)$$

where $\rho_\varepsilon = \int_{\mathbb{R}^d} g_\varepsilon M dv$, $u_\varepsilon = \int_{\mathbb{R}^d} g_\varepsilon v M dv$, $\theta_\varepsilon = \int_{\mathbb{R}^d} g_\varepsilon \left(\frac{v^2}{d} - 1 \right) M dv$

$$\Rightarrow \begin{cases} \partial_t \rho_\varepsilon + \frac{1}{\varepsilon} \operatorname{div} u_\varepsilon = 0 \\ \partial_t u_\varepsilon + \frac{1}{\varepsilon} \nabla(\rho_\varepsilon + \theta_\varepsilon) = -\frac{1}{\varepsilon} \operatorname{div} \int_{\mathbb{R}^d} g_\varepsilon \phi M dv \\ \frac{v^2}{2} \partial_t (\rho_\varepsilon + \theta_\varepsilon) + \frac{d+2}{2\varepsilon} \operatorname{div} u_\varepsilon = -\frac{1}{\varepsilon} \operatorname{div} \int_{\mathbb{R}^d} g_\varepsilon \psi M dv. \end{cases}$$

where $\phi(v) = v \otimes v - \frac{|v|^2}{d} \text{Id}$ and $\psi(v) = \left(\frac{v^2}{2} - \frac{d+2}{2} \right) v$.

(Note that $\varepsilon \rightarrow 0 \Rightarrow \frac{d}{dt} \int_{\mathbb{T}^d} (P+\theta) dx = 0 \Rightarrow \boxed{P+\theta=0}$ if true initially.)

Observe that $\phi, \psi \in \text{Ran } L$ (because $\perp \text{Ker } L$).

$\Rightarrow \exists \tilde{\phi}, \tilde{\psi} \in L^2(\text{Hdv})$ such that $L\tilde{\phi} = \phi$ and $L\tilde{\psi} = \psi$.

$$\Rightarrow \begin{cases} \partial_t u_\varepsilon + \frac{1}{\varepsilon} \nabla(P_\varepsilon + \theta_\varepsilon) = -\frac{1}{\varepsilon} \text{div} \int_{\mathbb{R}^d} (Lg_\varepsilon) \tilde{\phi} \text{Hdv} \\ \partial_t \left(\frac{d}{2} \theta_\varepsilon - P_\varepsilon \right) = -\frac{1}{\varepsilon} \text{div} \int_{\mathbb{R}^d} (Lg_\varepsilon) \tilde{\psi} \text{Hdv} \end{cases}$$

Use equation for g_ε to write $\frac{1}{\varepsilon} Lg_\varepsilon = \varepsilon^{m-1} Q(g_\varepsilon) - \varepsilon^{q-1} (\varepsilon \partial_t g_\varepsilon + v \cdot \nabla_x g_\varepsilon)$

From now on, suppose $m=1, q=1$ and $d=3$.

$$\Rightarrow \lim_{\varepsilon \rightarrow 0} \frac{1}{\varepsilon} L(g_\varepsilon) = Q(g) - v \cdot \nabla_x g = \frac{1}{2} L(g^2) - \text{div}(vg)$$

$$\stackrel{\substack{\uparrow \\ \text{careful} \\ \text{computation}}}{=} \frac{1}{2} u^t L(\phi) u + \theta u \cdot L(\psi) + \frac{1}{2} \theta^2 L\left(\frac{|v|^4}{4}\right) - \text{div}\left((P+\theta)v + \frac{v^3}{3} u + \phi u + \theta \psi\right)$$

$$= \frac{1}{2} u^t L(\phi) u + \theta u \cdot L(\psi) + \frac{1}{2} \theta^2 L\left(\frac{|v|^4}{4}\right) - \text{div}(\phi u + \theta \psi)$$

It can be shown that $\tilde{\phi}(v) = \alpha(|v|) \phi(v)$ and $\tilde{\psi}(v) = \beta(|v|) \psi(v)$.

(symmetry argument)

$$\Rightarrow \lim_{\varepsilon \rightarrow 0} \frac{1}{\varepsilon} \int_{\mathbb{R}^3} (Lg_\varepsilon) \tilde{\phi} \text{Hdv} = u \otimes u - \frac{|u|^2}{3} \text{Id} - \mu (\nabla_x u + \nabla_x^t u)$$

$$\text{and } \lim_{\varepsilon \rightarrow 0} \frac{1}{\varepsilon} \int_{\mathbb{R}^3} (Lg_\varepsilon) \tilde{\psi} \text{Hdv} = \frac{5}{2} \theta u - \frac{5}{2} k \nabla_x \theta$$

$$\text{with } \mu = \frac{1}{10} \int_{\mathbb{R}^3} \phi : \tilde{\phi} \text{Hdv} \text{ and } k = \frac{2}{15} \int_{\mathbb{R}^3} \psi \cdot \tilde{\psi} \text{Hdv}.$$

Let $P: L^2(dx) \rightarrow L^2(dx)$ denote the Leray projector onto divergence free vector fields, i.e., $P = \frac{\text{rot rot}}{-\Delta} = \text{Id} + \frac{\nabla \text{div}}{-\Delta}$.

$$\Rightarrow \begin{cases} \partial_t u + P(u \cdot \nabla_x u) - \mu \Delta_x u = 0 & , \quad \text{div } u = 0 \\ \partial_t \theta + u \cdot \nabla_x \theta - \kappa \Delta_x \theta = 0 & , \quad P + \theta = 0 \end{cases}$$

The incompressible Navier-Stokes-Fourier system.

Theorem: (Bardos, Golse, Levermore, Lions, Masmandi, Saint-Raymond, A.; 1991-2012)

Consider: • A physically relevant collision kernel which derives from an inverse power potential. For example: $b(v-u, \sigma) = |v-u|^\gamma b_0(\cos \theta)$, $-3 < \gamma \leq 1$
 $b_0(\cos \theta) \sim \frac{1}{\theta^{2+\nu}}$, $0 < \nu < 2$.

- Renormalized solutions $f_\varepsilon = \Pi(1 + \varepsilon g_\varepsilon) \in L_t^\infty L_{x,v}^1$ of $(\partial_t + v \cdot \nabla_x) g_\varepsilon = -\frac{1}{\varepsilon} \mathcal{L}(g_\varepsilon) + Q(g_\varepsilon)$ with well-prepared initial data.

Then, as $\varepsilon \rightarrow 0$:

- g_ε is weakly relatively compact in $L_{loc}^1(dt; L^1((1+v^2)\Pi dx dv))$
- each limit point g of g_ε satisfies $g = v \cdot u + \left(\frac{v^2}{2} - \frac{d+2}{2}\right) \theta$ where $(u, \theta) \in L_t^\infty L_x^2 \cap L_t^2 \dot{H}_x^1$ is a Leray solution of the incompressible Navier-Stokes-Fourier system.

Mathematical difficulties

1) Lack of a priori estimates! We only have the entropy inequality.

2) Renormalized solutions: $g_\varepsilon \in L^1_{loc}$ (at most $L \log L$) such that

$$(\varepsilon \partial_t + v \cdot \nabla_x) \frac{g_\varepsilon}{2 + \varepsilon g_\varepsilon} = \frac{-2}{\varepsilon(2 + \varepsilon g_\varepsilon)^2} L(g_\varepsilon) + \frac{2}{(2 + \varepsilon g_\varepsilon)^2} Q(g_\varepsilon)$$

3) Macroscopic conservation laws are not known to hold for renormalized solutions.

4) Time compactness: there are oscillations! (Use compensated compactness to filter out oscillations.)

5) Space compactness: there is some compactness thanks to velocity averaging lemmas. Not enough! Moreover, we need the nonlinear weak compactness estimate:

$$\frac{g_\varepsilon^2}{2 + \varepsilon g_\varepsilon} \text{ is weakly compact in } L^1_{loc}. \\ \text{(control concentrations)}$$