

Rigidity of non-compact static domains in hyperbolic space via positive mass theorems

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Outline of the talk

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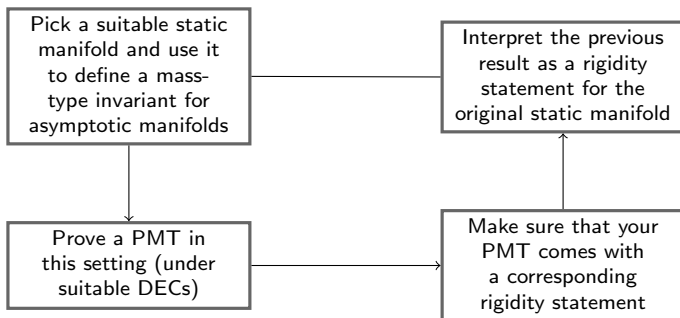
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Let g be a Riemannian metric in \mathbb{R}^n such that

- ▶ its **scalar curvature** R_g is non-negative everywhere;
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$$|g - \delta|_\delta = O_2(r^{-\sigma}), \quad \sigma > n - 2,$$

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¹R. Souam, Mean curvature rigidity of horospheres, hyperspheres, and [hyperplanes](#), *Archiv der Mathematik*, 2021.

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$$\mathcal{A} : \tilde{g} \mapsto \int_{\tilde{M}} (R_{\tilde{g}} - 2\tilde{\Lambda}) dv_{\tilde{g}} + 2 \int_{\partial\tilde{M}} (H_{\tilde{g}} - \tilde{\lambda}) d\sigma_{\tilde{g}},$$

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We say that a bordered Riemannian manifold $(M, g, \partial M)$ is a **static manifold with boundary**, with the pair $(\tilde{\Lambda}, \tilde{\lambda})$ as **cosmological constants**, if there exists $V \neq 0$ such that the equations above are satisfied. In this case, each such V is termed a **static potential**.

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- ▶ (\mathbb{R}_+^n, δ) is a static manifold with $(\tilde{\Lambda}, \tilde{\lambda}) = (0, 0)$. The corresponding PMT has been proved in [\[Almaraz, Barbosa, —, 2016\]](#) and its rigidity statement leads to (a sharper version of) the extrinsic rigidity result mentioned earlier.

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Equidistant hypersurfaces in \mathbb{H}^n

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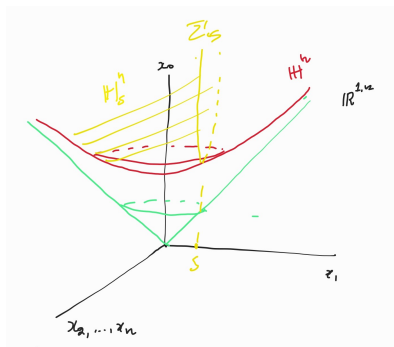
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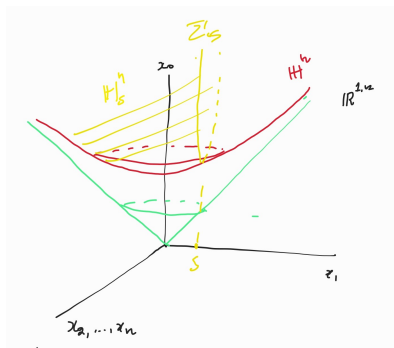
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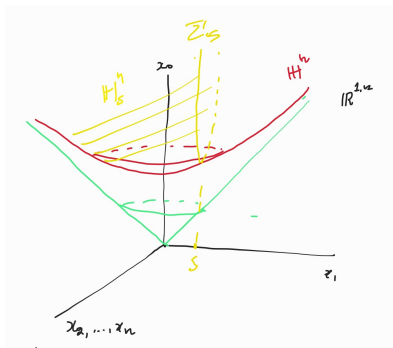
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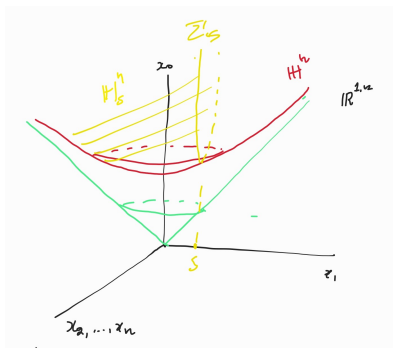
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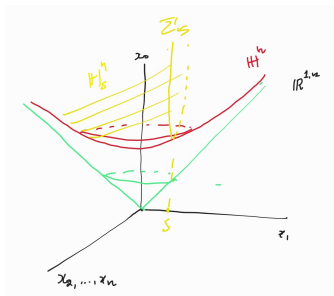
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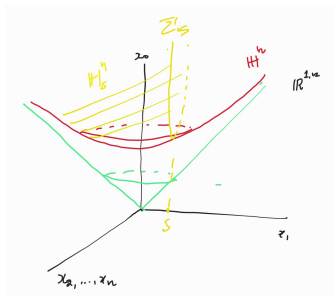
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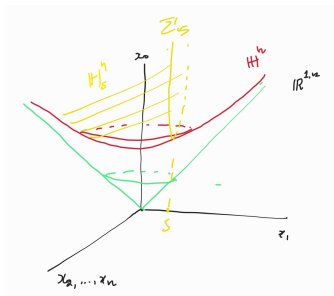
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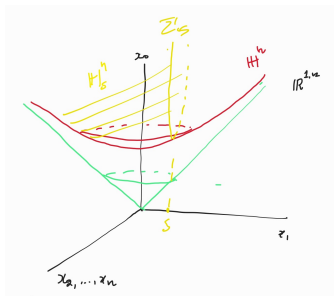
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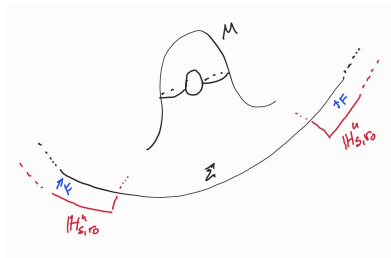
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Asymptotically hyperbolic manifolds modeled on \mathbb{H}_S^n

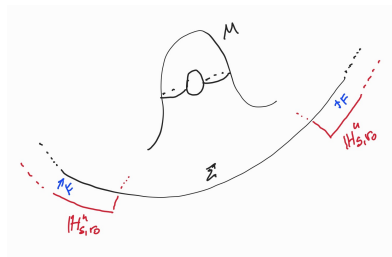
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We now consider the class of manifolds we will be interested in. Given $s \in \mathbb{R}$ as above, let us set $\mathbb{H}_{s,r_0}^n = \{x \in \mathbb{H}_s^n; r(x) \geq r_0\}$ for all r_0 large enough.



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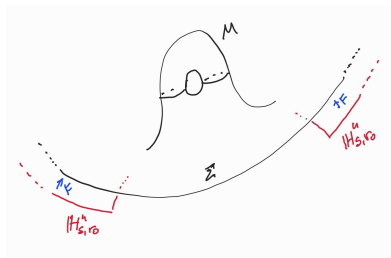
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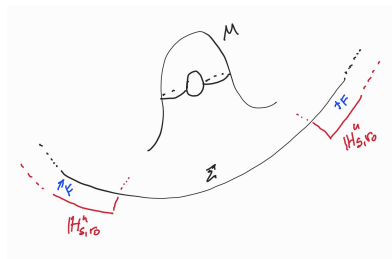
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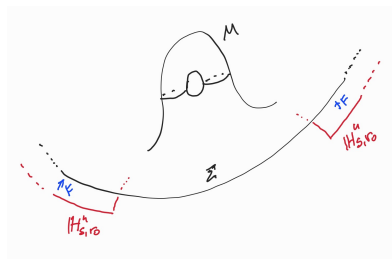
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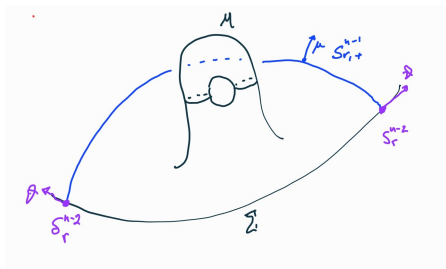
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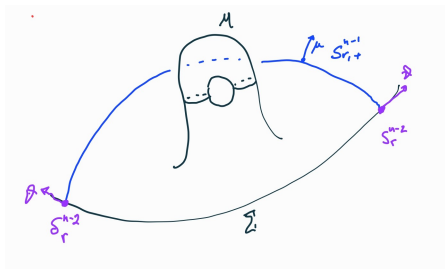
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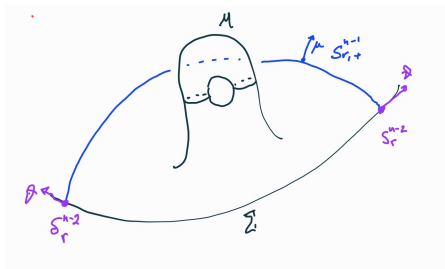
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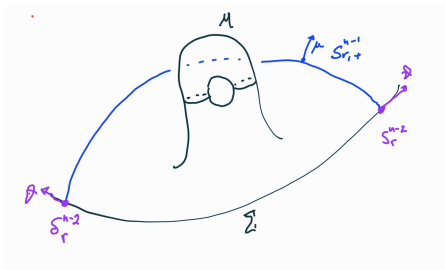
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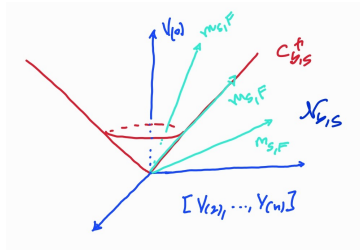
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It follows that ρ^s acts **isometrically** on $(\mathcal{N}_{b,s}, \langle \cdot, \cdot \rangle_{n-1,1}^s)$. In particular, the causal properties of $m_{s,F}$ are **chart independent**.



Invariance of the mass vector

- ▶ If F_1 and F_2 are charts at infinity then $F_{12} := F_1^{-1} \circ F_2 : \mathbb{H}_s^n \rightarrow \mathbb{H}_s^n$ satisfies $F_{12}^* b = b + O_2(r^{-\sigma})$. Because \mathbb{H}_s^n is "rigid" at infinity, there exists $A \in O^\uparrow(1, n-1)$ such that $F_{12} = A + O_2(r^{-\sigma})$

Proposition (equivariance of the mass)

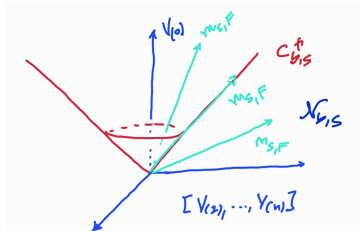
Under the conditions above,

$$m_{s,F_1} = \rho_{A^{-1}}^{s*}(m_{s,F_2}),$$

where ρ^{s*} is the dual representation.

- ▶ We may identify $\mathcal{N}_{b,s} \cong \mathbb{R}^{n-1,1}$ by introducing the "Lorentzian" metric $\langle \cdot, \cdot \rangle_{n-1,1}^s$ and declaring that $V_{(0)}$ is "time-like" (in the sense that $\langle V_{(0)}, V_{(0)} \rangle_{n-1,1}^s = 1$) and $\langle V_{(a)}, V_{(b)} \rangle_{n-1,1}^s = -\delta_{a,b}$, $2 \leq a, b \leq n$.

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Let (M, g, Σ) be an s -AH *spin* manifold with $R_g \geq -n(n-1)$ and $H_g \geq (n-1)\lambda_s$. Then, for any chart F as above, the mass vector $\mathfrak{m}_{s,F}$ is time-like and future directed unless it vanishes, in which case (M, g, Σ) is isometric to $(\mathbb{H}_s^n, b, \Sigma_s)$.

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$$m_s := \sqrt{\langle \mathfrak{m}_{s,F}, \mathfrak{m}_{s,F} \rangle_{1,n}^s},$$

does not depend on the chosen chart and may be regarded as the **total mass** of the isolated gravitational system whose (time-symmetric) initial data set is (M, g, Σ) . Hence, $m_s \geq 0$ with the equality holding if and only if (M, g, Σ) is isometric to $(\mathbb{H}_s^n, b, \Sigma_s)$.

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where c is Clifford multiplication, and the corresponding Killing-Dirac operators by $D^\pm : \Gamma(\mathbb{S}M) \rightarrow \Gamma(\mathbb{S}M)$ by the composition

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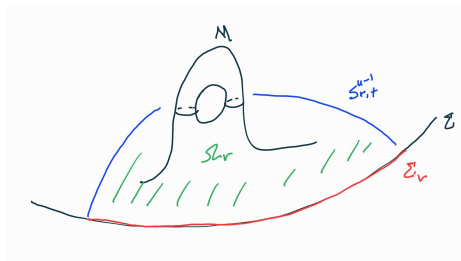
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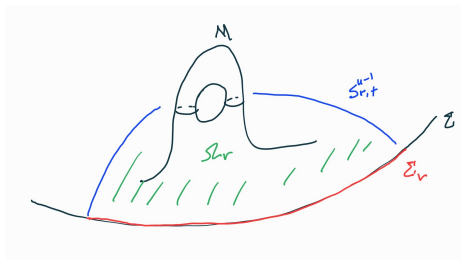
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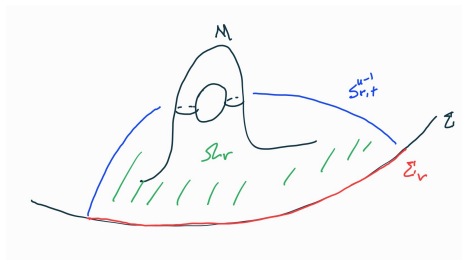
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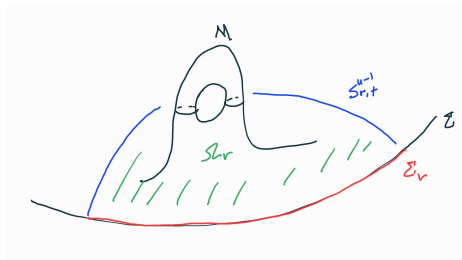
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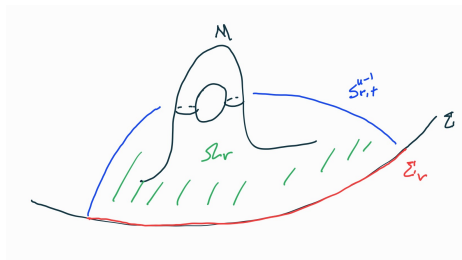
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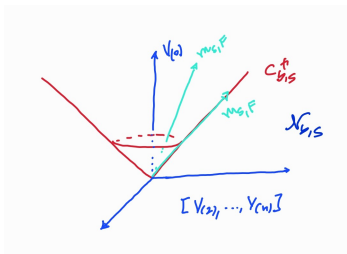
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The DECS then imply that $\langle m_{s,F}, V \rangle_{n,1} \geq 0$, for any $V \in \mathcal{C}_{b,s}^\uparrow$, which means that $m_{s,F}$ is time-like unless there exists a Killing spinor $\Psi^\theta \neq 0$ on M meeting the corresponding θ -boundary condition along Σ .



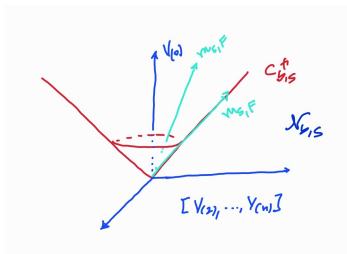
► The existence of Ψ^θ implies that g is Einstein ($\text{Ric}_g = -(n-1)g$) and Σ is totally umbilical (with $H_g = (n-1)\lambda_s$). In particular, $\Sigma \hookrightarrow M$ has the same second fundamental form as $\Sigma_s \hookrightarrow \mathbb{H}_s^n$.

How the mass formula implies our main result I

► Since any $V \in \mathcal{C}_{b,s}^\uparrow$ is of the form $V = V_\Phi$ for some Killing spinor Φ on \mathbb{H}_s^n , we have seen that, for any such V ,

$$\begin{aligned} \frac{1}{4} \langle m_{s,F}, V \rangle_{n,1} &= \int_M \left(|\nabla^\pm \Psi|^2 + \frac{R_g + n(n-1)}{4} |\Psi|^2 \right) dM \\ &\quad + \frac{1}{2} \int_\Sigma (H_g - (n-1)\lambda_s) |\Psi|^2 d\Sigma, \quad \Psi = \Psi_\Phi. \end{aligned}$$

The DECS then imply that $\langle m_{s,F}, V \rangle_{n,1} \geq 0$, for any $V \in \mathcal{C}_{b,s}^\uparrow$, which means that $m_{s,F}$ is time-like unless there exists a Killing spinor $\Psi^\theta \neq 0$ on M meeting the corresponding θ -boundary condition along Σ .



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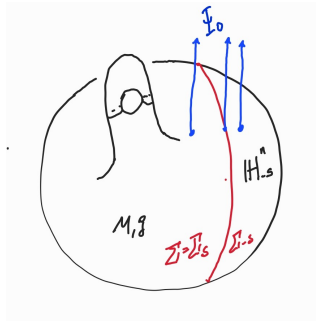
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This allows us to glue (M, g, Σ) to $(\mathbb{H}_{-s}^n, b, \Sigma_{-s})$ along the common boundary $\Sigma = \Sigma_{-s}$ to obtain a boundaryless n -manifold which is AH (with \mathbb{H}^n as its model at infinity), Einstein and carries a Killing spinor Ψ^θ . We conclude that this glued manifold is isometric to (\mathbb{H}^n, b) and hence (M, g, Σ) is isometric to $(\mathbb{H}_s^n, b, \Sigma_s)$, as desired.



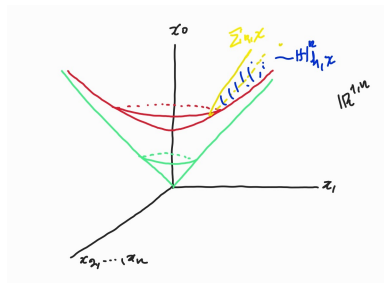
The horospherical case

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By means of the hyperboloid model $\mathbb{H}^n \hookrightarrow \mathbb{R}^{1,n}$, we consider the horoball

$$\mathbb{H}_h^n = \{x \in \mathbb{H}^n; x_0 - x_1 \leq 1\}.$$

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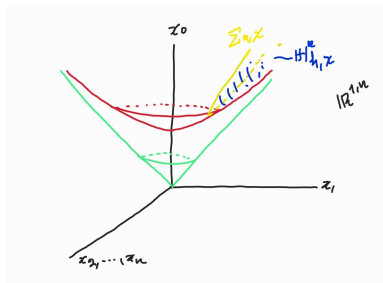


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Proposition

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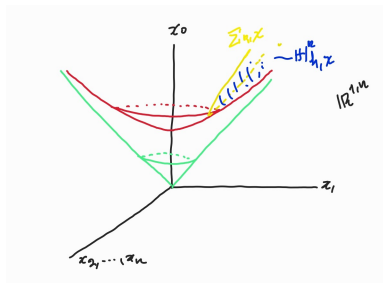
$$\mathcal{N}_{b,h} = [V_h, V_{(2)}, \dots, V_{(n)}], \quad V_h = (x_0 - x_1)|_{\mathbb{H}_h^n}.$$

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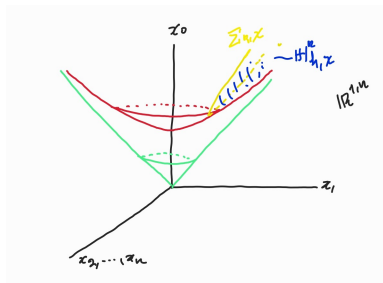
- Recall that the isometry group of \mathbb{H}_h^n may be identified to $O(n - 1) \times \mathbb{R}^{n-1}$, the group of euclidean motions of \mathbb{R}^{n-1} . Thus, we obtain a natural representation ρ^h of $O(n - 1) \times \mathbb{R}^{n-1}$ on $\mathcal{N}_{b,h}$ given by $\rho_A^h(V) = V \circ A^{-1}$.

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