



Holographic complexity

Javier MARTÍN GARCÍA

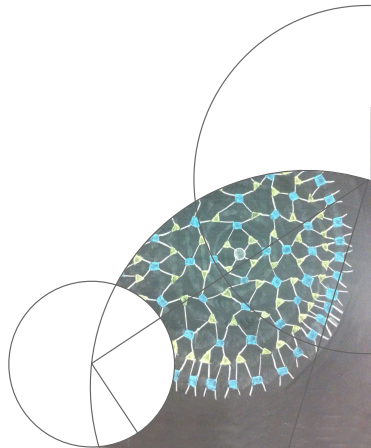
Based on:

J. L. F. BARBÓN & J. M-G. 1510.00349

& Ongoing work...



Instituto de
Física
Teórica
UAM-CSIC



① Introduction

The holographic dictionary

Quantum Information

② Quantum entanglement and computational complexity

Entanglement

Entanglement in the spacetime picture

Computational complexity

Complexity in the spacetime picture

③ Complexity of topological AdS black holes

General features

Non-extremal regime

Near-extremal regime

④ The Action/complexity proposal

From volume to action

Action for hyperbolic black holes

Comparison with C/V duality

⑤ Conclusions



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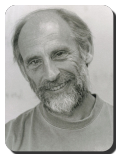
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Holography

- Quantum gravity in $d+1$ dimensions must be described by a non gravitational theory in d dimensions.
 - ★ Translation between both theories should be possible:
DICTIONARY
 - ★ Not so easy to find precise examples of holographic theories.
- Succesfull example: AdS/CFT
 - ★ Finding the entries of the dictionary is manageable.

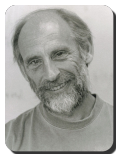


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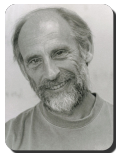


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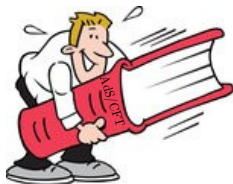
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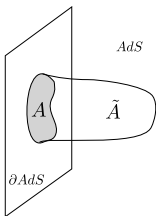
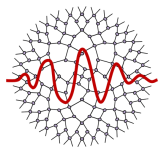
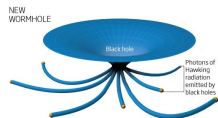
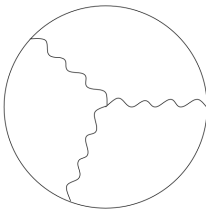
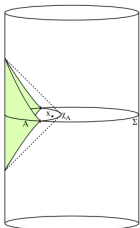
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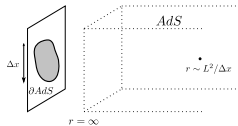
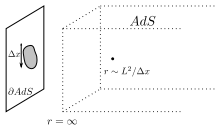
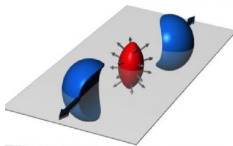
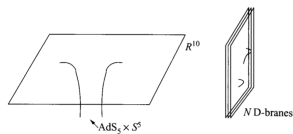
★ 1997:



Today:



$$\Delta = \frac{d}{2} + \sqrt{\frac{d^2}{4} + m^2 L^2}$$





DICTIONARY:
CHAPTER QI

AdS	CFT
	Entanglement?
	Complexity?

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Quantum entanglement

- Entanglement in QM

- ★ Failure of making the decomposition $|\psi\rangle = |\psi_A\rangle \otimes |\psi_B\rangle$
- ★ Quantified by entanglement entropy $S_A = -\text{Tr}_A[\rho_A \log \rho_A]$
- ★ Characterization of operator correlations

$$|\psi\rangle = |\psi_A\rangle \otimes |\psi_B\rangle \quad \Leftrightarrow \quad \langle \mathcal{O}_A \mathcal{O}_B \rangle - \langle \mathcal{O}_A \rangle \langle \mathcal{O}_B \rangle = 0.$$

- Tensor Network

- ★ 2-Qubit entangled state $|\Psi\rangle = \sum_{i,j} \alpha_{ij} |\psi_i\rangle_A |\psi_j\rangle_B$.
- ★ General entangled state $|\Psi\rangle = \sum_{\mu_1, \dots, \mu_n} c_{\mu_1, \dots, \mu_n} |\psi_{\mu_1}\rangle_1 \dots |\psi_{\mu_n}\rangle_n$

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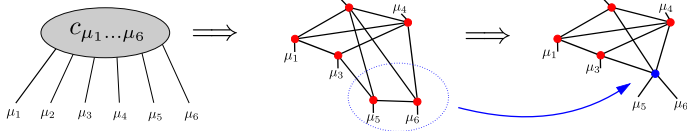
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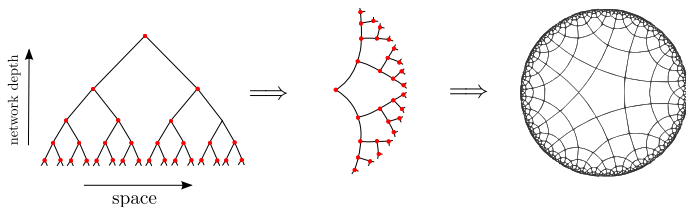
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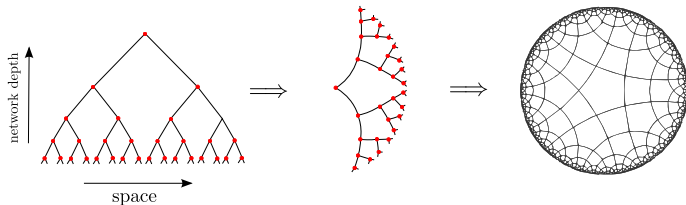
Entanglement in the spacetime picture

- Tensor network for a CFT
 - ★ Scale invariance \implies Tree-like structure TN
 - ★ New (non-physical) dimension: Network depth
 - ★ Mimics a discretized hyperbolic space
- Holographic entanglement entropy
 - ★ Entanglement between two regions \iff number of links cut to separate the TN in two.
 - ★ Ryu-Takayanagi formula $S_A = \frac{\text{Area}(\tilde{A})}{4G}$.



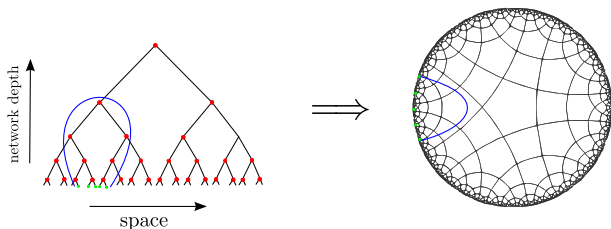
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- Entanglement and spacetime connectivity
 - ★ Eternal AdS black hole $|\Psi\rangle = \sum_i e^{-\beta E_i/2} |E_i\rangle \otimes |E_i\rangle \implies$ Two entangled BH connected by a wormhole
 - ★ Classically connected spacetime \iff Superposition of disconnected ones
 - ★ ER=EPR: Any entangled system is connected by a wormhole

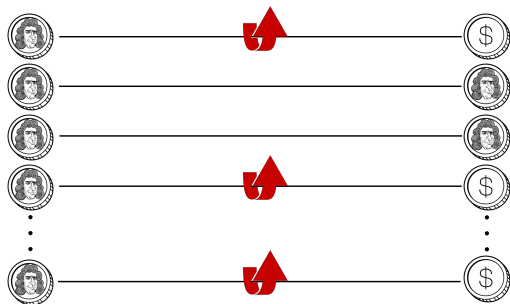


DICTIONARY:
CHAPTER QI

AdS	CFT
Connectivity	Entanglement
Minimal surfaces	S_A
ER	EPR
	Complexity?

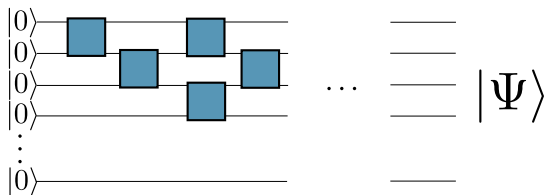
Computational complexity

- Computer science: Minimum number of 'simple' operations needed to get a state from a 'simple' reference one. Example: Given a string of K bits, get some state 011010.. from 000000...



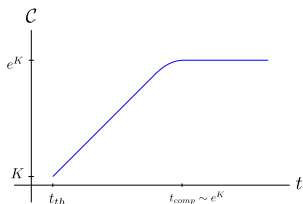
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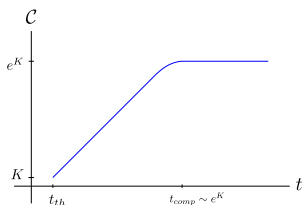
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- Growth of complexity
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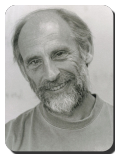


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 - ★ $\frac{dC}{dt} \propto TS$

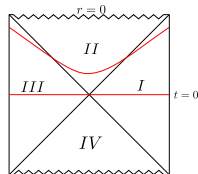
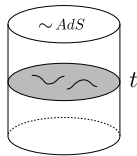
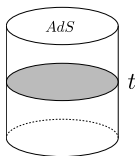


Complexity in the spacetime picture

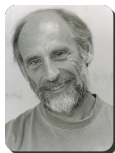


L. Susskind

- Tensor networks
 - ★ Complex states have bigger tensor networks
 - ★ Complexity \Leftrightarrow size of the tensor network
- Volume/complexity relation: The complexity of $|\psi(t)\rangle$ is proportional to the volume of a maximal slice in the dual spacetime that passes through t .
- Black holes
 - ★ Growth of quantum complexity is encoded as the growth of the Einstein-Rosen Bridge

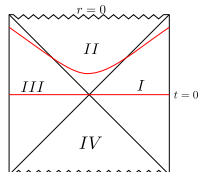
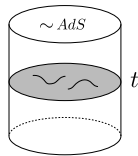
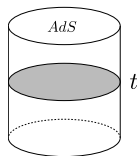


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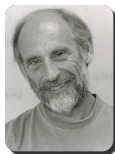


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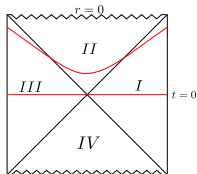
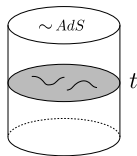
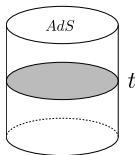


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Maximal volumes	Complexity

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General features

- Topological black holes

- ★ Metric

$$ds^2 = -f(r)dt^2 + f(r)^{-1}dr^2 + \frac{r^2}{l^2}dH_{d-1}^2$$

$$f(r) = -1 + \frac{r^2}{l^2} - \frac{\mu}{r^{d-2}}$$

- ★ Dual to two CFT's on a hyperboloid [Emparan]

- ★ Degenerate system: $\lim_{T \rightarrow 0} S \neq 0$

- Slicing conditions

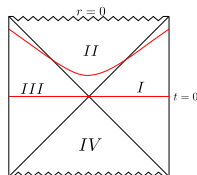
- ★ Spacelike Cauchy surfaces

- ★ 'Nice' slices: Stay away from singularities

- ★ Asymptotically match constant t surfaces far away

- ★ Foliation of the entire exterior region

- ★ Asymptote to constant r_m surface in the interior for long times.



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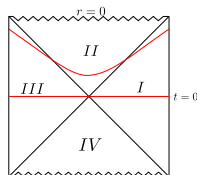
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Non-extremal regime

- Features

- ★ $T_H \gg 1$
- ★ One horizon. Schwarzschild-Ads-like topology.

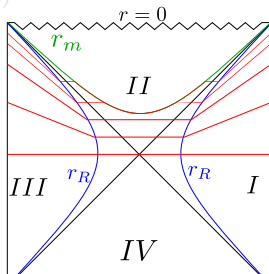
- Metric Patching

- ★ Exterior ($r \gg 1$) $\implies f_E(r) = -1 + r^2 \implies$ Constant t surface
- ★ Rindler region ($r \sim r_h$) $\implies f_R(r) = 4\pi T_H(r - r_h) \implies$ Horizontal planes (in X,T coordinates)
- ★ Interior ($r \ll r_h$)

- Results

- ★ Exterior: Constant contribution
- ★ Wormhole:

$$V \sim \sqrt{\frac{d}{1 - 2^{-1/d}}} G_N S T_H t$$



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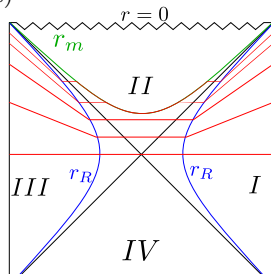
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$$V \sim \sqrt{\frac{d}{1 - 2^{-1/d}}} G_N S T_H t$$



Non-extremal regime

- Features

- ★ $T_H \gg 1$
- ★ One horizon. Schwarzschild-Ads-like topology.

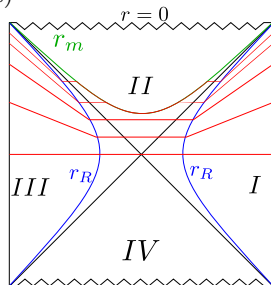
- Metric Patching

- ★ Exterior ($r \gg 1$) $\implies f_E(r) = -1 + r^2 \implies$ Constant t surface
- ★ Rindler region ($r \sim r_h$) $\implies f_R(r) = 4\pi T_H(r - r_h) \implies$ Horizontal planes (in X,T coordinates)
- ★ Interior ($r \ll r_h$)

- Results

- ★ Exterior: Constant contribution
- ★ Wormhole:

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Near-extremal regime

- Features

- ★ $T_H < \frac{1}{2\pi}$
- ★ Two horizons r_{\pm} . Timelike singularity. Degenerate system.

- Metric Patching

- ★ Exterior ($r \gg 1$) \implies Constant t surface
- ★ $AdS_{1+1} \times H^{d-1}$ region ($r_h \ll r \ll 1$) \implies Constant t surface
- ★ Rindler region ($r \sim r_h$) \implies Horizontal planes (in X,T coordinates)
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- Results

- ★ Exterior: Constant contribution
- ★ $AdS_{1+1} \times H^{d-1}$ region: Constant divergent contribution $\sim \log T^{-1}$
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① Introduction

The holographic dictionary

Quantum Information

② Quantum entanglement and computational complexity

Entanglement

Entanglement in the spacetime picture

Computational complexity

Complexity in the spacetime picture

③ Complexity of topological AdS black holes

General features

Non-extremal regime

Near-extremal regime

④ The Action/complexity proposal

From volume to action

Action for hyperbolic black holes

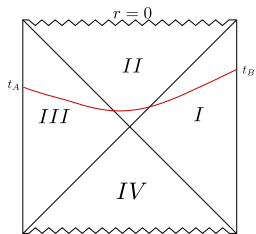
Comparison with C/V duality

⑤ Conclusions



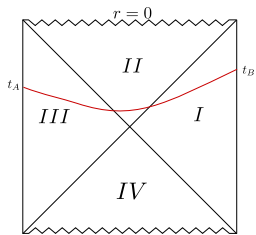
From volume to action

- Unpleasant features of volume/complexity duality
 - ★ Arbitrary scale $\mathcal{C} \sim \frac{V}{Gl}$, $l \sim \ell_{AdS}, r_h, \dots$
 - ★ Why should the maximal slice play a preferred role?
- A higher dimensional object might solve the problem
 - ★ $\mathcal{C} \sim \frac{V\ell_{AdS}}{Gl_{AdS}^2} \sim \frac{W}{Gl_{AdS}^2} \sim \frac{\Lambda}{G} \int \sqrt{g} dV \sim A$
 - ★ Consider all possible foliations



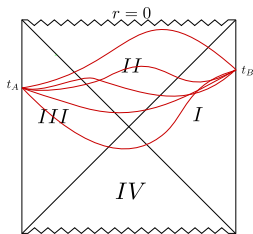
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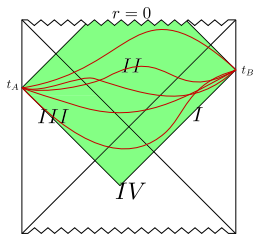
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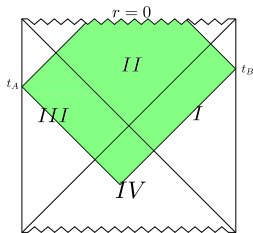
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The Action/complexity proposal

- The complexity of $|\psi(t)\rangle$ is given by the on-shell action of its gravitational dual evaluated on the WdW patch corresponding to t .

$$\mathcal{C} = \frac{\mathcal{A}}{\pi\hbar}$$

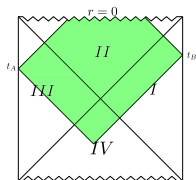
[Brown, Roberts, Susskind, Swingle, Zhao]

- Some features

- ★ Absence of arbitrary scales
- ★ No preferred foliation
- ★ Recovers the nice features of V/C
 - BH complexity growth: $\frac{d\mathcal{C}}{dt} = 2M \sim ST$
 - Shockwave tests [Susskind, Stanford]
- ★ Connection to Lloyd's bound

- Drawbacks

- ★ The YGH term is ill-defined for null surfaces and joints. New prescriptions for these quantities still ambiguous and *ad hoc*. [Lehner, Myers, Poisson, Sorkin...]



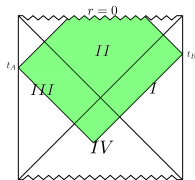
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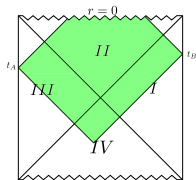
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A curiosity

- 'Extended' black hole thermodynamics

$$\text{Smarr formula: } (d-2)M = (d-1)TS - 2PV$$

- Action growth for spherical black holes factorizes in these terms
[Coach, Fischler, Nguyen]

$$\star \delta S_{\text{bulk}} \sim PV\delta t$$

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DICTIONARY:
CHAPTER QI

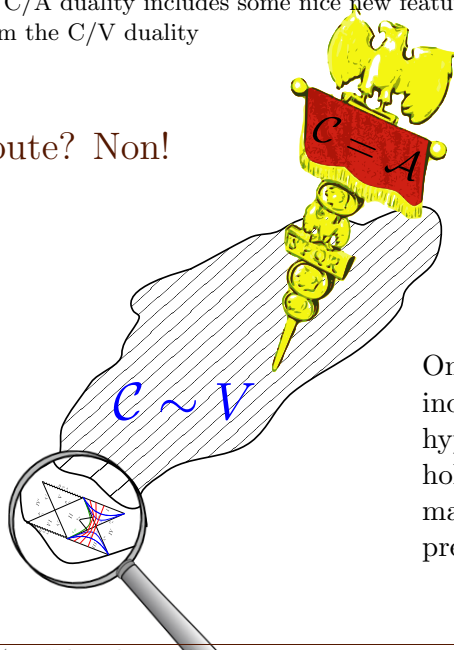
AdS	CFT
Connectivity	Entanglement
Minimal surfaces	S_A
ER	EPR
Maximal volumes	Complexity_a
WDW action	Complexity_b

It seems that the C/A duality includes some nice new features and recovers all the results from the C/V duality



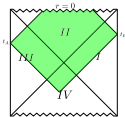
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Toute? Non!



One small set of indomitable hyperbolic black holes still resist to match both prescriptions

WdW action for hyperbolic black holes



- Complexity for cold BH's is **finite** [Myers, Chapman, Marrochio]

- Complexity growth

- * Hot black holes: $\frac{d\mathcal{C}}{dt} = 2M_{AdS}$

- * Cold black holes do not compute! $\frac{d\mathcal{C}}{dt} = 0$

- For the cold ones 'thermodynamic factorization' no longer holds

- * $\delta S_{\text{bulk}} \sim (PV + r_-^d)\delta t$

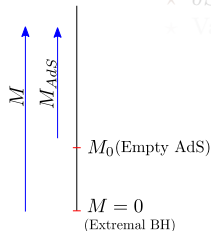
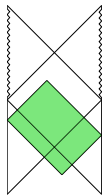
- * $\delta S_{\text{bound.}} = 0$

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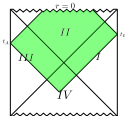
- * Vacua ambiguities

- * $\frac{dS}{dt} = TS - \frac{2PV_+}{d-1} - \frac{(d-2)}{(d-1)}M_{AdS}$

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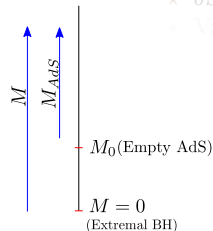
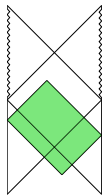
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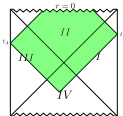
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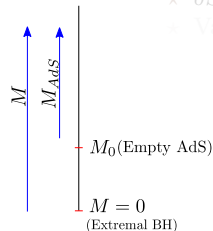
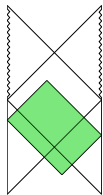
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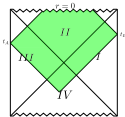
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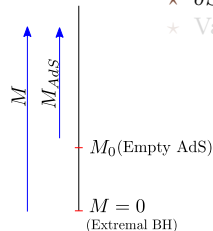
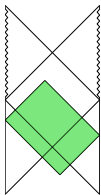
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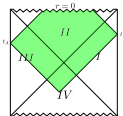
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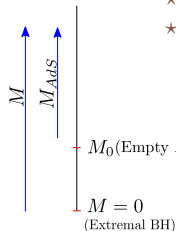
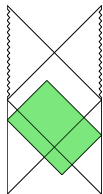
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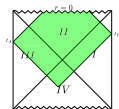
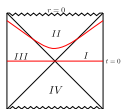
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Comparison between both prescriptions



General features	Complexity/Volume	Complexity/Action
Definition	Preferred slicing	Covariant definition
Scale	Ambiguous	Fixed
TN interpretation	# of tensors	?
$d\mathcal{C}/dt$ in BH	$\sim ST$	$2M$
Technicalities	Optimization	Boundary terms
Cold hyp. BH		
\mathcal{C} in AdS_{1+1} throat	$\sim \log T$	Finite
$d\mathcal{C}/dt$	$\sim S_0 T$	0

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Conclusions & Outlook

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Questions



Muito obrigado!

