Black Hole Collapse in CFT

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BH Collapse in CFT



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[arXiv: 1603.04856 & to appear]





Fonds national suisse Schweizerischer Nationalfonds Fondo nazionale svizzero Swiss National Science Foundation Fidelity Decay vs. Unitarity (featuring: the Iberians)



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Laura Garcia-Alvarez & QUTIS

[arXiv: 1607.08560 & 1701.XXYYZ]







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"introduction"



black holes are thermodynamic systems

their entropy is proportional to the area of the event horizon

information loss paradox:

a BH formed from a pure state will evolve into a mixed state (of Hawking radiation) holography:

a theory of quantum gravity should have information ~ area

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holography:

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General Plan

AdS/CFT relates gravity (often in AdS) to unitary field theory (often CFT)

Lots of progress gravity - CFT (my favorite: AdS/CMT)

Less known about CFT → (quantum) gravity

→ despite developments in CFT, CMT:

- time evolution and spread of entanglement
- thermalization of closed quantum systems (e.g. via eigenstates)
- non-perturbative methods (e.g. bootstrap)

Thermalization \rightarrow BH formation (& evaporation)

Outline

1. Introduction

2. Part I: Unitarity vs. Gravity

3. Part II: The anti-information paradox

4. Conclusions

Part I: Unitarity vs. Gravity

"the trouble with black holes"

 outgoing Hawking radiation is thermal ρ_{Gibbs}(T_H)

- a horizon ${\mathcal H}$ cloaks the singularity

- initial pure state $|\Psi\rangle$ of matter collapses inwards



black holes evaporate



The Paradox

- gravity as an EFT implies pure to mixed evolution
- fundamentally incompatible with a unitary S-matrix
- 1. quantum gravity is non-unitary
- 2. gravity EFT makes no sense
- 3. (subtle) corrections to Hawking result





hence AdS/CFT only allows for options 2 & 3.



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the anti-information loss paradox:

how does an obviously unitary theory lose information?

(of a first-principles calculation in holographic CFT)

1. define an initial state in CFT which forms a black hole

2. understand time evolution in strong-coupling regime

3. diagnose signs of information loss & recovery

4. translate this into a consistent picture of bulk quantum gravity

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- 4. translate this into a consistent picture of bulk quantum gravity
 → wouldn't that be nice?

"unitarity constraints"

Unitarity vs Thermalization

(constraints on long-time correlations from unitarity)

Correlations in a closed quantum system, e.g.

$$G(t) = \mathrm{tr}\rho\mathcal{O}(t)\mathcal{O}(0)$$

Time average over a large time T cannot vanish by unitarity

$$\lim_{T \to \infty} \overline{|G(t)|^2} \neq 0$$

Need to assume spectrum is generic (no specific ordering principle)

→ fails for integrable theories
→ connection with ETH

Unitarity vs Thermalization

(constraints on long-time correlations from unitarity)

$$\rho = e^{-\beta H}$$



see also [Barbon & Rabonivici]

Spectral Probes

(to appear with del Campo and Molina-Vilaplana)

Essence of 2-point function: dephasing at late times

$$G(t) = \frac{1}{Z} \sum_{i,j} e^{-\beta E_i - it(E_i - E_j)} |O_{ij}|^2$$

Conveniently captured in spectral form factor

$$g(\beta, t) = \frac{1}{Z^2} |Z(\beta + it)|^2$$

Well-studied object in random matrix theory

Recently re-surfaced in context of SYK [Cotler et al.] & 2D CFT [Dyer & Gur-Ari]



initial decay (~information loss), then "dip, ramp & plateau"

can we characterise these properties more generally?

Fidelity Decay

(a new perspective on spectral form factor)

Prepare the system in the "thermofield double" state on $\mathcal{H}^{\otimes 2}$

$$|\psi(\beta)\rangle := \frac{1}{\sqrt{Z}} \sum_{n} e^{-\frac{\beta}{2}E_n} |n\rangle \otimes |n\rangle$$

Consider time evolution, U(t,0), with respect to $H=H_{\rm s}\otimes {\bf 1}$

We can re-express the spectral form factor as the overlap

$$\mathcal{F}(t) = |\langle \psi(\beta) | U(t,0) | \psi(\beta), 0 \rangle|^2$$

 \Rightarrow Unitarity constraints on behaviour of "fidelity" $\mathcal{F}(t)$ (e.g QSL)

Selected Results [del Campo, Molina-Vilaplana, JS]

1. Initial Gaussian decay governed by "Zeno Time" $\tau_Z = \langle \Delta E \rangle_\beta$



2. Intermediate decay slower than exponential

e.g. SYK & RMT $\sim 1/t^3$

3. Very late time: non-commuting limits & information loss

c.f. rest of talk!

Upshot

QFT correlations probe information loss

Unitarity demands non-trivial late-time behaviour $\mathcal{O}\left(e^{-S}\right)$ (assume finite-size system)

Fidelity of thermofield double state allows to map bounds on spectral form factor to QSL

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Part II: the antiinformation loss paradox

Looking for the Right Place

0D matrix models(IOP,...): connection to geometry? 2D black hole (CGHS): solvable but very different SYK/2D black hole: Einstein dual,...?

3D story shares salient features of 4D (and higher) in fact central to micro-state counting success (D1-D5)

the trouble: no local degrees of freedom (Achucarro & Townsend):

$$S_{3D} = S_{CS}[A] - S_{CS}[\bar{A}]$$

other side of the coin: CFT₂ puts powerful tools at our disposal also see [Fitzpatrick, Kaplan,...]

3D Gravity + Matter

→ add matter: get local dof. BUT need new tools

focus on a universal sector, by defining a 1/c expansion:

→ any microscopic theory in this class defines some 3D quantum gravity theory (sparse spectrum)

3D gravity + matter non-trivial, but solvable → ideal place to study BH puzzles!

From bulk point of view this is $1/G_N$ expansion

CFT₂ gives a non-perturbative definition of quantum gravity.

The Black Hole in the Tin Can



global AdS_{d+1}

Throw in a shell of n dust particles

smooth limit: $n \to \infty$

BH collapse: Vaidya metric

Use light operators ${\mathcal Q}$ to probe geometry as function of t

remark: certain quantities such as entanglement entropy are sensitive to behind horizon physics (away from equilibrium)

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Translating to the CFT



Interrogating the CFT

Start probing the physics via 2n + p correlations

$$G(1,2,\ldots p) = \langle \mathcal{V} | \mathcal{Q}_1, \ldots \mathcal{Q}_p | \mathcal{V} \rangle$$

we want to approach smooth, semi-classical gravity

$$c \to \infty$$
$$n \to \infty$$
$$\sigma \to 0$$
$$E \sim nh_{\psi}/\sigma \to \mathcal{O}(c)$$

infinite-point correlations in strongly-coupled CFT!

Benefits of 2D CFT

in the semi-classical limit (large c), get sum of exponentials

$$G(1, 2, \dots p) = \sum_{\text{blocks}} a_k e^{-\frac{c}{6}f_k^{(n)}(1, 2, \dots p)}$$

correlator approximated by largest term, the identity block

"it from id"

the dominant contribution comes from the identity Virasoro block, that is the unit operator **id** and all its descendants *T*, ∂*T*, *T*² *T*∂*T*..., (multi-graviton exchange in bulk)

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subleading corrections exponentially suppressed in e^{-c} ~ e^{-1/G}

still need to calculate the semi-classical block:

CONFORMAL SCALAR FIELD ON THE HYPERELLIPTIC CURVE AND CRITICAL ASHKIN-TELLER MULTIPOINT CORRELATION FUNCTIONS

Al.B. ZAMOLODCHIKOV

Scientific Council of "Cybernetics", Academy of Sciences, USSR

Received 3 December 1986

A multipoint conformal block of Ramond states of the two-dimensional free scalar field is calculated. This function is related to the free energy of the scalar field on the hyperelliptic Riemann surface under a particular choice of boundary conditions. Being compactified on the still need to calculate the semi-classical block:

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USSR: fighting hyper-intelligent Robot overlords with CFT?

The Monodromy Method

each contraction of operators in the plane defines a cycle



fix monodromies of y''(z) + Ty(z) = 0 $\xrightarrow{c_k} f_k^{(n)}(1, 2, \dots p)$

Taking the smooth Limit

generally a hard problem, big simplification occurs for n $\longrightarrow \infty$



stress tensor \mapsto distribution on $\Sigma_{(\infty)}$

continuum monodromy method $f_0^\infty(1,2,\ldots,p)$



3D semi-classical gravity $\mathcal{L}_{ ext{geo}}(1,2,\ldots p)$

Two-point Autocorrelation

let us now return to the black hole and compute

 $G(t) = \mathrm{tr}\rho\mathcal{O}(t)\mathcal{O}(0)$



can be done analytically:

$$G(t_1, t_2) = \left(\frac{1}{\pi T} \cos\left(\frac{t_1}{2}\right) \sinh\left(\pi T t_2\right) - 2\sin\left(\frac{t_1}{2}\right) \cosh\left(\pi T t_2\right)\right)^{-2\Delta^{\mathcal{Q}}}$$

Physical Consequences

not (yet) known from gravity (but matches known limits) \implies CFT prediction for 3D gravity

The correlation function decays without bound at large time

$$G(t_1, t_2) \sim \exp(-\frac{2\pi\Delta^{\mathcal{Q}}t}{\beta})$$

Manifestly in conflict with unitarity: **CFT loses information!**

Can also compute entanglement entropy of interval A

$$S(A) \to S_{\text{Gibbs}}(A;T) \longrightarrow \rho(A) = \rho_{\text{Gibbs}}(A;T)$$

Restoring Unitarity

This is the anti-information paradox: what happened to unitarity?

$$|G(t)| = \left| \sum_{n,k} e^{i(E_n - E_k)t} \Psi_n^*(\mathcal{V}) \langle n | \mathcal{Q} | k \rangle \langle k | \mathcal{Q} | \mathcal{V} \rangle \right| \neq 0$$

ightarrow correlations cannot become arbitrarily small in $|\mathcal{V}\rangle$

Neglected contributions exponentially suppressed at t=0 (must be present due to crossing symmetry)

$$\sum_{k \neq \text{vac}} a_k e^{-\frac{c}{6} f_k^{\infty}(1,2,\dots,p)} \sim e^{-S}$$

restore unitary at large time \rightarrow non-perturbative effects in 1/G_N

Conclusions

time-dependent 3D quantum gravity with matter in 1/c expansion 'it from id' \rightarrow ideal arena to think about quantum BHs

unitarity constraints on 2-point functions, spectral form factor,... can be mapped to fidelity decay and quantum speed limits (QSL) → bounds on scrambling exponent?

CFT correlation functions seemingly violate unitarity (naïve). non-perturbative corrections in c restore unitarity

on gravity side these correspond to non-perturbative effects in $G_{N.}$ geometric interpretation? bulk interpretation?

are the corrections universal? make a black hole that evaporates,...

thank you!

entanglement entropy

Q-type operators \rightarrow twist insertions: $G_q(t) = \langle \mathcal{V} | \sigma_q(t, \ell_1) \tilde{\sigma}_q(t, \ell_2) | \mathcal{V} \rangle$

 z_1

$$S(A) = \lim_{q \to 1} \frac{1}{1 - q} G_q(t)$$

crossing points $z_{c1} \& z_{c2} \leftrightarrow$ refraction at bulk shell

it from id → require trivial monodromy on smile contour

write
$$z_1=e^{i heta_1}, z_2=e^{i(heta_1+L)}$$
 & continue to Lorentzian time $heta_1=t$

maximize S(A) over crossing points \rightarrow parametric equation for S(t)

 z_2

 \dot{z}_{c2}

entanglement entropy

Implicit formula for growth of entanglement entropy:

$$t = \frac{\beta}{2\pi} \cosh^{-1} \left\{ \cosh\left(2\pi Tq\right) + 2\pi T \tan\left(\frac{L}{2} - q\right) \sinh\left(2\pi Tq\right) \right\}$$
$$S_{EE} = \frac{c}{3} \log \left\{ \frac{\sin\left(\frac{L}{2} - q\right) \cosh\left(2\pi Tq\right) + \frac{1}{2\pi T} \left[1 + \frac{1}{2} \left\{1 + 4\pi^2 T^2\right\} \tan^2\left(\frac{L}{2} - q\right)\right] \cos\left(\frac{L}{2} - q\right) \sinh\left(2\pi Tq\right)}{\epsilon_{UV}/2} \right\}$$

matches **exactly** global AdS₃ Vaidya:

- thermal at late time
- EE growth = change of channel
- sees beyond horizon



CFT calculation shows that purity of state is preserved:

alternative picture: IN-IN computation

