

Holographic entanglement entropy for perturbative higher-curvature gravities

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[2012.14033 - P. Bueno, J. Camps, AVL]

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Why entanglement entropy?

Entanglement entropy allows us to gain significant physical insight into quantum field theories.

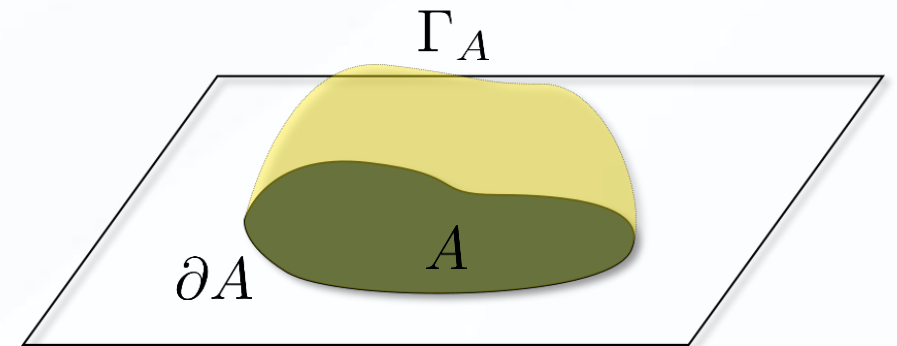
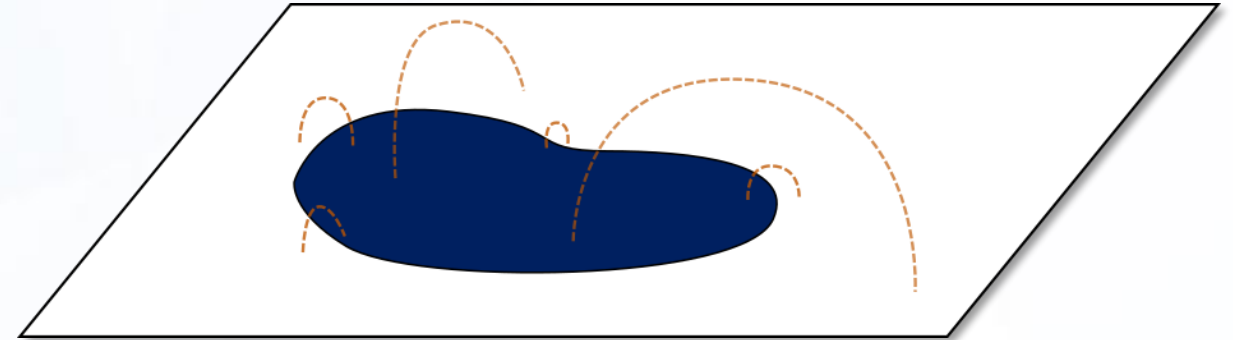
But it is usually difficult to calculate!

[0603001, Ryu, Takayanagi]

In a holographic theory dual to GR, it is possible to compute entanglement entropy via a geometric problem (minimization of the area of a bulk surface).

$$S_{\text{HEE}}^{\text{E}}(A) = \frac{\mathcal{A}(\Gamma_A)}{4G_N}$$

Is this a **deep statement about quantum gravity**?



Why higher-curvature gravity?

Higher-curvature gravity theories appear naturally within the context of AdS/CFT:

- ➔ They are expected as **quantum or stringy corrections** to the classical bulk action, corresponding to corrections to the strong coupling and large- N limits of the CFT.
- ➔ Their holographic duals are inequivalent to the one defined by Einstein gravity (*e.g.*, $a \neq c$ in 4d). This has been very fruitful in discovering universal properties of CFTs, or as toy models to probe interesting CFT physics.

We aim to understand **how entanglement entropy is to be calculated** in the bulk when we include **higher-curvature corrections**.

Why perturbative?

The general form of the entanglement entropy functional is known:

[1310,5713, X. Dong]

[1310,6659, J. Camps]

$$S_{\text{HEE}}^{\mathcal{L}(\text{Riem})}(A) = S_{\text{Wald}} + S_{\text{Anomaly}}$$

$$S_{\text{Wald}} = -2\pi \int_{\Gamma_A} d^{d-1}y \sqrt{h} \epsilon_{\mu\nu} \epsilon_{\rho\sigma} \frac{\partial \mathcal{L}_E}{\partial R_{\mu\nu\rho\sigma}}$$

$$S_{\text{Anomaly}} = 2\pi \int_{\Gamma_A} d^{d-1}y \sqrt{h} \sum_{\alpha} \left(\frac{\delta \partial^2 \mathcal{L}_E}{\partial \text{Riem}^2} K^2 \right)_{\alpha} \frac{1}{1 + q_{\alpha}}$$

$$\left(\frac{\delta \partial^2 \mathcal{L}_E}{\partial \text{Riem}^2} K^2 \right) \equiv 2 \left[\perp^{\lambda_1 \lambda_2} (\perp_{\mu_1 \mu_2} \perp_{\nu_1 \nu_2} - \epsilon_{\mu_1 \mu_2} \epsilon_{\nu_1 \nu_2}) + \epsilon^{\lambda_1 \lambda_2} (\perp_{\mu_1 \mu_2} \epsilon_{\nu_1 \nu_2} + \epsilon_{\mu_1 \mu_2} \perp_{\nu_1 \nu_2}) \right] \\ \times \frac{\partial^2 \mathcal{L}_E}{\partial R_{\mu_1 \rho_1 \nu_1 \sigma_1} \partial R_{\mu_2 \rho_2 \nu_2 \sigma_2}} K_{\lambda_1 \rho_1 \sigma_1} K_{\lambda_2 \rho_2 \sigma_2}$$

Just a projected second derivative contracted with K's!

$$\epsilon_{\mu\nu} \equiv \epsilon_{ab} n^a_{\mu} n^b_{\nu}, \quad \perp_{\mu\nu} \equiv \delta_{ab} n^a_{\mu} n^b_{\nu}, \quad g_{\mu\nu} \equiv h_{\mu\nu} + \perp_{\mu\nu}$$

The holographic surface should be obtained by extremizing the new functional. If we work **perturbatively, it can be taken to be the RT surface.**

Why perturbative?

$$S_{\text{HEE}}^{\mathcal{L}(\text{Riem})}(A) = S_{\text{Wald}} + S_{\text{Anomaly}}$$

$$S_{\text{Anomaly}} = -2\pi \int_{\Gamma_A} d^{d-1}y \sqrt{h} \sum_{\alpha} \left(\frac{\delta \partial^2 \mathcal{L}_E}{\delta \text{Riem}^2} K^2 \right)_{\alpha} \frac{1}{1+q_{\alpha}}$$

The functional comes from applying a replica trick-like technique in the bulk which involves **regularizing a conical singularity** at the entanglement surface:

[1304,4926, Lewkowycz, Maldacena]

➔ Different Riemann tensor components contribute differently at the conical singularity, thus they must be expanded as:

$$R_{z\bar{z}z\bar{z}} = \tilde{R}_{z\bar{z}z\bar{z}} - \frac{1}{8} \underbrace{K^{aij} K_{aij}}_{q_{\alpha} = 1}, \quad \underbrace{R_{z\bar{z}zi}}_{q_{\alpha} = 1/2}, \quad R_{z\bar{z}ij} = \tilde{R}_{z\bar{z}ij} - \underbrace{2K_{z[i}{}^k K_{\bar{z}]j}{}_k}_{q_{\alpha} = 1}, \quad \dots \quad \begin{array}{l} \epsilon_{z\bar{z}} = i/2 \\ \perp_{z\bar{z}} = 1/2 \end{array}$$

➔ The **regularization is not unique** (*splitting problem*), and it must be fixed by the EoM of the theory. This has only been done for GR, and perturbative corrections to it.

[Exceptions: Lovelock, f(R), quadratic gravity]

[1503,05538, R.-X. Miao]

[1605,08588, J. Camps]

The anomaly expansion

$$S_{\text{Anomaly}} = -2\pi \int_{\Gamma_A} d^{d-1}y \sqrt{h} \sum_{\alpha} \left(\frac{8\partial^2 \mathcal{L}_E}{\partial \text{Riem}^2} K^2 \right)_{\alpha} \frac{1}{1+q_{\alpha}}$$

This is a mess (and difficult to understand what it is doing):

$$\mathcal{L}_E = R_{\mu\nu}{}^{\rho\sigma} R_{\rho\sigma}{}^{\delta\gamma} R_{\delta\gamma}{}^{\mu\nu} \quad \left(\frac{8\partial^2 \mathcal{L}_E}{\partial \text{Riem}^2} K^2 \right) = 48 R^{zi\bar{z}j} K_{zi}{}^k K_{\bar{z}jk} \quad [K_a = 0]$$

$$\sum_{\alpha} (R^{zi\bar{z}j})_{\alpha} = \tilde{R}^{zi\bar{z}j} - \underbrace{K^{zik} K_{\bar{z}j}{}^k}_{q_{\alpha} = 1}$$

$$\sum_{\alpha} (R^{zi\bar{z}j})_{\alpha} \frac{1}{1+q_{\alpha}} = \tilde{R}^{zi\bar{z}j} - \frac{1}{2} K^{zik} K_{\bar{z}j}{}^k = R^{zi\bar{z}j} + \frac{1}{2} K^{zik} K_{\bar{z}j}{}^k$$

$$\sum_{\alpha} \left(\frac{8\partial^2 \mathcal{L}_E}{\partial \text{Riem}^2} K^2 \right)_{\alpha} \frac{1}{1+q_{\alpha}} = 48 R^{zi\bar{z}j} K_{zi}{}^k K_{\bar{z}jk} + 24 K^{zik} K_{\bar{z}j}{}^k K_{zi}{}^l K_{\bar{z}jl}$$

Generates higher-order terms in K!

The anomaly expansion revisited

It is possible to understand this anomaly term in a simpler way:

$$S_{\text{Anomaly}} = 2\pi \int_{\Gamma_A} d^{d-1}y \sqrt{h} \sum_{S=0}^{\infty} \frac{1}{S!} \int_0^1 du 2u : \left(-(1-u^2)\hat{\partial}_A - (1-u)\hat{\partial}_B \right)^S : \left(\frac{8\partial^2 \mathcal{L}_E}{\partial \text{Riem}^2} K^2 \right)$$

$$\hat{\partial}_A = \left[\frac{1}{2} \perp^{\lambda_1 \lambda_2} h^{\tau_1 \tau_2} h^{\omega_1 \omega_2} \epsilon_{\mu\nu} \epsilon_{\rho\sigma} - 2\epsilon^{\lambda_1 \lambda_2} h_{\rho}^{\tau_1} h_{\sigma}^{\tau_2} h^{\omega_1 \omega_2} \epsilon_{\mu\nu} - 2 \perp^{\lambda_1 \lambda_2} h_{\mu}^{\tau_1} h_{\rho}^{\omega_1} h_{\nu}^{\tau_2} h_{\sigma}^{\omega_2} \right. \\ \left. - 2 \left(\perp^{\lambda_1 \lambda_2} \perp_{\mu\rho} + \epsilon^{\lambda_1 \lambda_2} \epsilon_{\mu\rho} \right) h_{\nu}^{\tau_1} h_{\sigma}^{\tau_2} h^{\omega_1 \omega_2} \right] K_{\lambda_1 \tau_1 \omega_1} K_{\lambda_2 \tau_2 \omega_2} \frac{\partial}{\partial R_{\mu\nu\rho\sigma}}$$

$$+ 2 \left(\perp_{\mu_2}^{\mu_1} \perp_{\rho_2}^{\rho_1} - \epsilon_{\mu_2}^{\mu_1} \epsilon_{\rho_2}^{\rho_1} \right) h_{\nu_2}^{\nu_1} h_{\sigma_2}^{\sigma_1} R_{\mu_1 \nu_1 \rho_1 \sigma_1} \frac{\partial}{\partial R_{\mu_2 \nu_2 \rho_2 \sigma_2}}$$

$$\hat{\partial}_B = 4 \left[\perp_{\mu_2}^{\mu_1} h_{\nu_2}^{\nu_1} h_{\rho_2}^{\rho_1} h_{\sigma_2}^{\sigma_1} + \perp_{\mu_2}^{\mu_1} \perp_{\nu_2}^{\nu_1} \perp_{\rho_2}^{\rho_1} h_{\sigma_2}^{\sigma_1} \right] R_{\mu_1 \nu_1 \rho_1 \sigma_1} \frac{\partial}{\partial R_{\mu_2 \nu_2 \rho_2 \sigma_2}}$$

Generates higher order terms in K (certain Riemann tensor components)

Only terms up to S=n-2 in an n-th order Lagrangian

This is both algorithmically and conceptually clearer!

The anomaly expansion: general results

- 1 Theories of the form $f(\text{Ricci})$ do not have anomaly term (perturbatively):

$$S_{\text{HEE}}^{\mathcal{L}(\text{Ricci})} = \frac{\mathcal{A}(\Gamma_A)}{4G} + \frac{\lambda}{8G_N} \int_{\Gamma_A} d^{d-1}y \sqrt{h} \frac{\partial \mathcal{L}}{\partial R_{\mu\nu}} \perp_{\mu\nu} + \mathcal{O}(\lambda^2)$$

$$I_E^{\mathcal{L}(\text{Ricci})} = -\frac{1}{16\pi G_N} \int d^{d+1}x \sqrt{|g|} \left[\frac{d(d-1)}{L^2} + R + \lambda \mathcal{L}(g_{\mu\nu}, R_{\rho\sigma}) \right]$$

- 2 Riccis do not generate K's. An n -th order Lagrangian containing n_R Riemann tensors and $n - n_R$ Riccis (or scalars) has extrinsic curvatures up to the power $2n_R - 2$

$$S_{\text{Anomaly}} \sim \int_{\Gamma_A} \sum \text{Ricci}^{n-n_R} \left(\text{Riem}^{n_R-2} K^2 + \text{Riem}^{n_R-3} K^4 + \dots + \text{Riem} K^{2n_R-4} + K^{2n_R-2} \right)$$

$$\mathcal{L} \sim \text{Riem}^{n_R} \text{Ricci}^{n-n_R}$$

In particular, densities with 0 or 1 Riemann tensors have no anomaly (perturbatively).

The anomaly expansion: general results

- 3 The anomaly term for Lovelock theories has a nice expression:

$$S_{\text{Anomaly}}^{\mathcal{X}_{2n}} = 2\pi \int_{\Gamma_A} d^{d-1}y \sqrt{h} \left[2K_{aik} K^a{}_{lj} \frac{\partial}{\partial R_{ijkl}} \right]^{-1} \left[\exp \left(2K_{aik} K^a{}_{jl} \frac{\partial}{\partial R_{ijkl}} \right) - 1 \right] \left(\frac{8\partial^2 \mathcal{X}_{2n}}{\partial \text{Riem}^2} K^2 \right)$$

$$\mathcal{X}_{2n}(R) \equiv \frac{1}{2^n} \delta^{\mu_1 \mu_2 \dots \mu_{2n-1} \mu_{2n}}_{\nu_1 \nu_2 \dots \nu_{2n-1} \nu_{2n}} R^{\nu_1 \nu_2}_{\mu_1 \mu_2} \dots R^{\nu_{2n-1} \nu_{2n}}_{\mu_{2n-1} \mu_{2n}}$$

Combined with the Wald term, this becomes totally intrinsic, and it coincides with the Jacobson-Myers functional:

$$S_{\text{HEE}}^{\mathcal{X}_{2n}} = -4\pi n \int_{\Gamma_A} d^{d-1}y \sqrt{h} \mathcal{X}_{2(n-1)}(\mathcal{R})$$

With this general form of the functional, **we have obtained the contribution of cubic (and quartic) theories:**

$$I_E^{\text{Riem}^3} = -\frac{1}{16\pi G_N} \int d^{d+1}x \sqrt{|g|} \left[\frac{d(d-1)}{L^2} + R + L^4 \sum_{i=1}^8 \beta_i \mathcal{L}_i^{(3)} \right] \left| \begin{array}{ll} \mathcal{L}_1^{(3)} \equiv R_{\mu\nu}{}^{\rho\sigma} R_{\rho\sigma}{}^{\delta\gamma} R_{\delta\gamma}{}^{\mu\nu} & \mathcal{L}_2^{(3)} \equiv R_{\mu\nu}{}^{\rho\sigma} R_{\rho\sigma}{}^{\delta\gamma} R_{\delta\gamma}{}^{\mu\nu} \\ \mathcal{L}_3^{(3)} \equiv R_{\mu\nu\rho\sigma} R^{\mu\nu\rho\sigma} R^{\sigma\delta} & \mathcal{L}_4^{(3)} \equiv R_{\mu\nu\rho\sigma} R^{\mu\nu\rho\sigma} R \\ \mathcal{L}_5^{(3)} \equiv R_{\mu\nu\rho\sigma} R^{\mu\rho} R^{\nu\sigma} & \mathcal{L}_6^{(3)} \equiv R_{\mu}{}^{\nu} R_{\nu}{}^{\rho} R_{\rho}{}^{\mu} \\ \mathcal{L}_7^{(3)} \equiv R_{\mu\nu} R^{\mu\nu} R & \mathcal{L}_8^{(3)} \equiv R^3 \end{array} \right.$$

$$S_{\text{HEE}}^{\text{Riem}^3} = \frac{\mathcal{A}(\Gamma_A)}{4G_N} + \frac{L^4}{4G_N} \int_{\Gamma_A} d^{d-1}y \sqrt{h} \sum_{i=1}^8 \beta_i \Delta_i^{(3)} + \mathcal{O}(\beta_i^2)$$

...

$$\Delta_2^{(3)} = + 3R^{ab\rho\sigma} R_{ab\rho\sigma} - 6K_{ai}{}^k K_{bjk} (R^{aibj} - R^{biaj}) - 6K_{aik} K^{ajk} R^{bi}{}_{bj} + 3K_{ai}{}^j K_{bj}{}^k K^a{}_{k}{}^l K^b{}_{l}{}^i - 6K_{ai}{}^j K^a{}_{j}{}^k K_{bk}{}^l K^b{}_{l}{}^i$$

...

$$\Delta_6^{(3)} = + \frac{3}{2} R^{a\mu} R_{a\mu}$$

$$S_{\text{HEE}}^{\mathcal{L}(\text{Riem})}(A) = S_{\text{Wald}} + S_{\text{Anomaly}}$$

$$S_{\text{Wald}} = -2\pi \int_{\Gamma_A} d^{d-1}y \sqrt{h} \epsilon_{\mu\nu} \epsilon_{\rho\sigma} \frac{\partial \mathcal{L}_E}{\partial R_{\mu\nu\rho\sigma}}$$

$$S_{\text{Anomaly}} = 4\pi \int_{\Gamma_A} d^{d-1}y \sqrt{h} \int_0^1 du u : e^{-F(u)} : \left(\frac{8\partial^2 \mathcal{L}_E}{\partial \text{Riem}^2} K^2 \right), \quad F(u) \equiv \underbrace{(1-u^2)\hat{\partial}_A + (1-u)\hat{\partial}_B}$$

Projected second derivative

Operators made out of
projected derivatives and K's

Universal terms of entanglement entropy

Entanglement entropy of a (smooth) region A in a d-dimensional CFT:

Typical lengthscale of A
or IR regulator

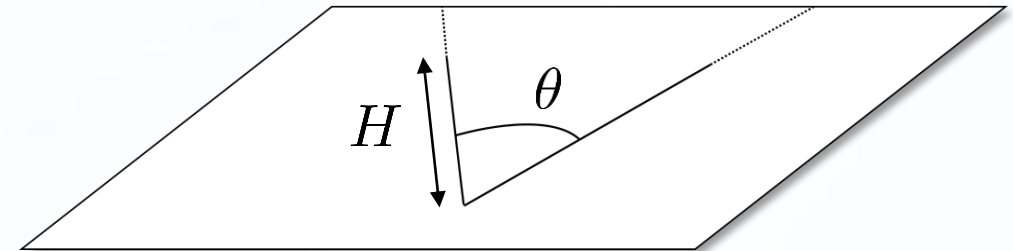
$$S_{\text{EE}}(A) = p_1 \left(\frac{H}{\delta}\right)^{d-2} + p_2 \left(\frac{H}{\delta}\right)^{d-4} \dots + \begin{cases} p_{(d-2)/2} (H/\delta)^2 + c \log(H/\delta) & [d \text{ even}] \\ p_{(d-1)/2} (H/\delta) + \tilde{c} & [d \text{ odd}] \end{cases}$$

UV regulator
Non-universal terms
Universal terms

Geometric singularities introduce new terms. For a corner in d=3 in vacuum:

$$S_{\text{EE}}(A) = b_1 \frac{H}{\delta} - a(\theta) \log(H/\delta) + b_0$$

↑
Universal piece
(corner function)



Holographic corner functions

Some well-known properties:

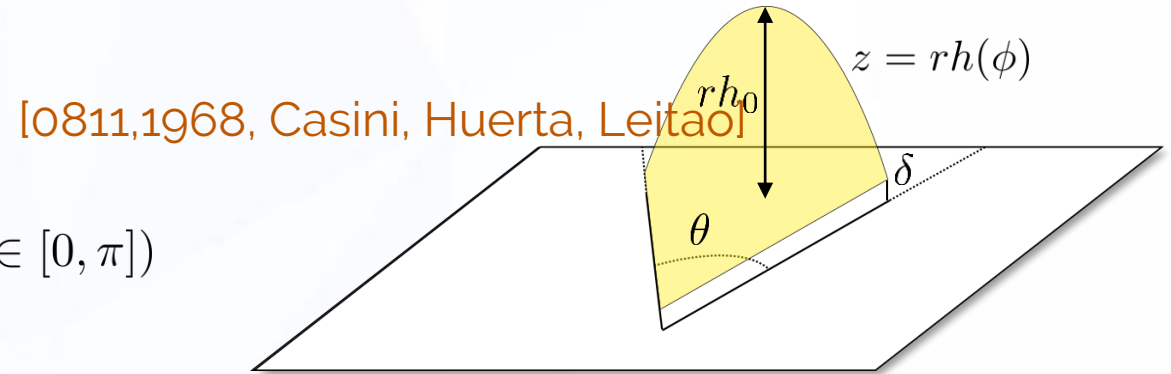
$$a(\theta) \geq 0, \quad \partial_\theta a(\theta) \leq 0, \quad \partial_\theta^2 a(\theta) \geq -\frac{\partial_\theta a(\theta)}{\sin \theta} \quad (\theta \in [0, \pi])$$

Sharp limit: $a(\theta) \underset{\theta \rightarrow 0}{\sim} \frac{\kappa}{\theta} + \mathcal{O}(\theta)$

Smooth limit: $a(\theta) \underset{\theta \rightarrow \pi}{\sim} \sigma (\theta - \pi)^2 + \mathcal{O}(\theta - \pi)^4$

Only explicitly known for simple models (free bosons/fermions, EMI model, ...) and, holographically, for Einstein gravity:

$$S_{\text{HEE}}^{\text{E}} = \frac{L_\star^2}{2G_N} \int_{\delta/h_0}^H \frac{dr}{r} \int_0^{\theta/2-\epsilon} d\phi \frac{\sqrt{1+h^2+\dot{h}^2}}{h^2}$$



[0811.1968, Casini, Huerta, Leitaol]

$$ds^2 = \frac{L_\star^2}{z^2} (d\tau^2 + dz^2 + dr^2 + r^2 d\phi^2)$$

$$a_{\text{E}}(\theta) = \frac{L_\star^2}{2G_N} \int_0^{+\infty} dy \left[1 - \sqrt{\frac{1+h_0^2(1+y^2)}{2+h_0^2(1+y^2)}} \right]$$

$$\theta = \int_0^{h_0} dh \frac{2\sqrt{1+h_0^2}h^2}{\sqrt{1+h^2} \sqrt{(h_0^2-h^2)(h_0^2+(1+h_0^2)h^2)}}$$

Holographic corner functions

Quadratic theories do not add anything new: $\mathcal{L}_1^{(2)} = R_{\mu\nu\rho\sigma}R^{\mu\nu\rho\sigma}$, $\mathcal{L}_2^{(2)} = R_{\mu\nu}R^{\mu\nu}$, $\mathcal{L}_3^{(2)} = R^2$

$$a_{\text{Riem}^2}(\theta) = [1 - 6\alpha_2 - 24\alpha_3] a_{\text{E}}(\theta)$$

But (some) cubic theories do!

$$\mathcal{L}_1^{(3)} \equiv R_{\mu}{}^{\rho}{}_{\nu}{}^{\sigma} R_{\rho}{}^{\delta}{}_{\sigma}{}^{\gamma} R_{\delta}{}^{\mu}{}_{\gamma}{}^{\nu} , \quad \mathcal{L}_2^{(3)} \equiv R_{\mu\nu}{}^{\rho\sigma} R_{\rho\sigma}{}^{\delta\gamma} R_{\delta\gamma}{}^{\mu\nu} , \quad \dots$$

$$a_{\text{Riem}^3}(\theta) = [1 + 6\beta_1 + 12\beta_2 + 6\beta_3 + 24\beta_4 + 27\beta_5 + 27\beta_6 + 108\beta_7 + 432\beta_8] a_{\text{E}}(\theta) + \sum_{i=1}^2 \beta_i g_i(\theta)$$

$$g_1(\theta) \equiv + \frac{L_*^2}{2G_N} \int_0^{+\infty} \frac{3(1+h_0^2) [3+h_0^2(5+4y^2) + 2h_0^4(1+y^2)^2]}{[1+h_0^2(1+y^2)]^{7/2} \sqrt{2+h_0^2(1+y^2)}} dy$$

$$g_2(\theta) \equiv - \frac{L_*^2}{2G_N} \int_0^{+\infty} \frac{6(1+h_0^2) [3+h_0^2(7+8y^2) + 4h_0^4(1+y^2)^2]}{[1+h_0^2(1+y^2)]^{7/2} \sqrt{2+h_0^2(1+y^2)}} dy$$

This is the **first holographic corner function** with a functional dependence **different from the Einstein gravity one!**

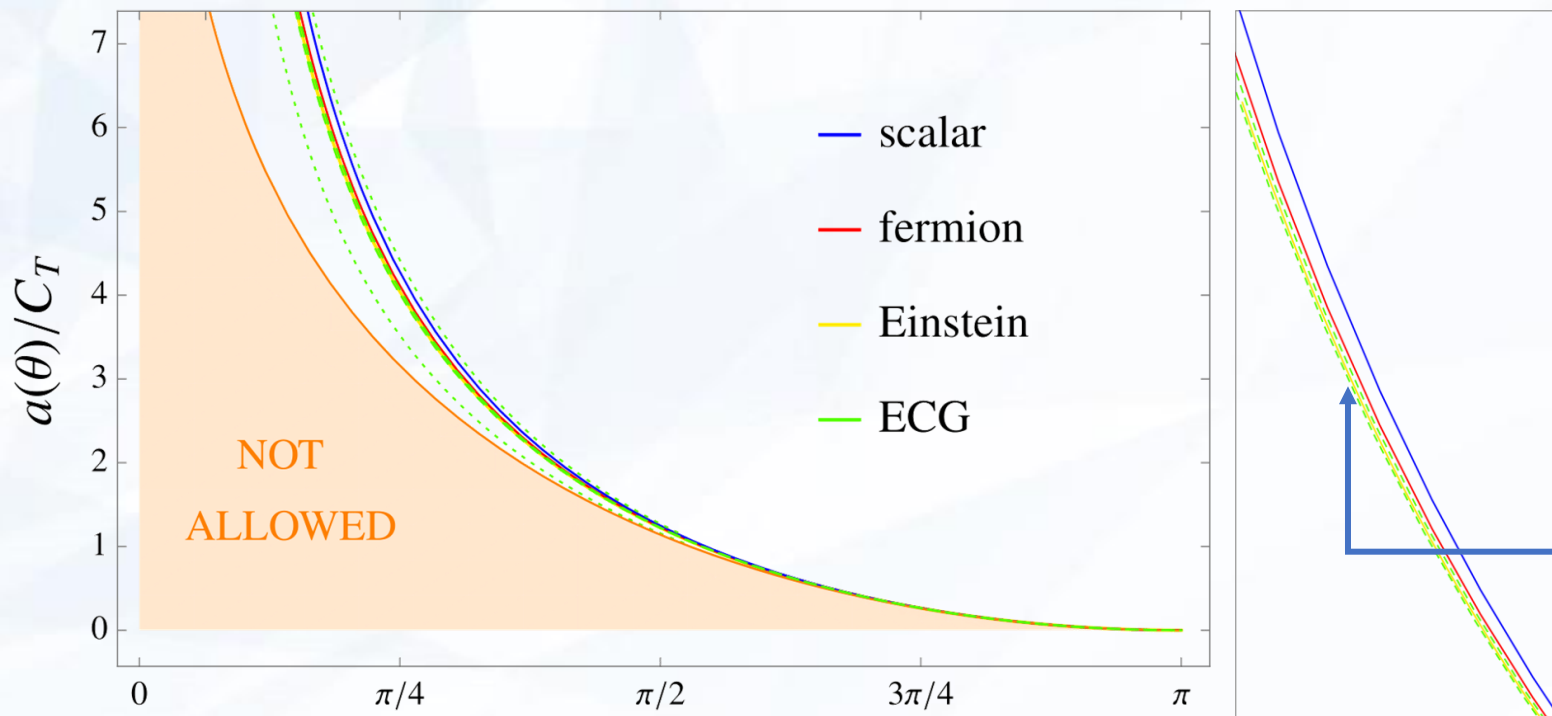
Holographic corner functions

Take Einsteinian Cubic Gravity as an example:

[1607,06463, P. Bueno, P. Cano]

$$I_E^{\text{ECG}} = -\frac{1}{16\pi G_N} \int d^4x \sqrt{|g|} \left[\frac{6}{L^2} + R - \frac{\mu_{\text{ECG}} L^4}{8} \mathcal{P} \right]$$

$$\left[\mathcal{P} \equiv 12\mathcal{L}_1^{(3)} + \mathcal{L}_2^{(3)} - 12\mathcal{L}_5^{(3)} + 8\mathcal{L}_6^{(3)} \right]$$





All extremely close when normalized by C_T !

[1505,04804, Bueno, Myers, Witczak-Krempal]

Some are below the Einstein gravity result, which was conjectured to be a lower bound

- ✓ We have presented a **rewriting of the entanglement entropy functional** for perturbative HCG, which is both **computationally and conceptually simpler**.
- ✓ This was used to compute the functional corresponding to cubic corrections – and quartic, see paper!
- ✓ It also gave us a better understanding of the structure of the functional: n_R Riemann tensors produce $2n_R - 2$ extrinsic curvatures.
- ✓ We used the cubic functionals to compute universal contributions to the EE. For a corner in $d=3$, we have obtained the **first holographic corner function which is not proportional to the Einstein gravity one** (for other shapes, see paper!).

Future work:

-  Can this be generalized for non-perturbative corrections?
-  Is the new formula useful in other contexts, *e.g.*, to prove second laws for BHs?

Obrigado!