#### JT gravity - a review

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#### Introduction: 2d dilaton gravity models

#### JT classical gravity

Classical solution Schwarzian quantum mechanics

#### JT quantum gravity: disk amplitudes

Partition function Boundary correlators and its gravitational physics

#### JT quantum gravity: higher topology Gravitational amplitudes JT gravity as a matrix integral

#### Some recent developments

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We need different coupling to matter: dilaton gravity

$$S = \frac{1}{16\pi G} \int d^2x \sqrt{-g} \left(\Phi R + V(\Phi)\right) + S_{bdy} + S_{matter} \Phi$$
 is dilaton field

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- Dimensional reduction (s-wave) of 3d pure  $\Lambda < 0$  gravity
- Appears as near-horizon theory of near-extremal higher-dimensional black holes
- Describes low-energy sector of all (known) SYK-like models
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Knowing  $R(x) \Rightarrow$  we know everything about the (local) geometry Geometry fixed as AdS<sub>2</sub>:  $ds^2 = \frac{-dF^2 + dZ^2}{Z^2}$ ,  $Z \ge 0$ Poincaré patch (frame) of AdS<sub>2</sub>, boundary at Z = 0

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Poincaré patch:  $ds^2 = -\frac{4dUdV}{(U-V)^2}$ Found in near-horizon regime of extremal black hole ▶ BH frame:  $U(u) = \tanh\left(\frac{\pi}{\beta}u\right), V(v) = \tanh\left(\frac{\pi}{\beta}v\right)$  $ds^2 = -\frac{\pi^2}{\beta^2} \frac{4}{\sinh^2(\frac{\pi}{2}(u-v))} du dv$ Found in near-horizon regime of near-extremal black hole Using radial coordinate  $r \sim \coth \frac{2\pi}{\beta}(u-v)$ :  $ds^{2} = -(r^{2} - r_{h}^{2})dt^{2} + \frac{dr^{2}}{r^{2} - r_{h}^{2}}$ is black hole solution with ADM mass  $E \sim r_h^2$ , and Hawking temperature  $T \sim r_h \longrightarrow$  first law  $T \sim \sqrt{E}$ 

# Important frames in $AdS_2$ (2)

Penrose diagram



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 $\Rightarrow S = -C \int dt \{F, t\}$ ,  $C = \frac{a}{16\pi G}, \{F, t\} = \frac{F''}{F'} - \frac{3}{2} \left(\frac{F''}{F'}\right)^2$ 

Almheiri-Polchinski '15, Jensen '16, Maldacena-Stanford-Yang '16, Engelsöy-TM-Verlinde '16 F(t) = time reparametrization in terms of proper time t

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Generalization to include coupling to matter (CFT for simplicity):  $-C\frac{d}{dt} \{F, t\} = \frac{dE}{dt} = T_{vv}(t) - T_{uu}(t)$ in terms of boundary values of matter energy Total energy is changed by injection and extraction from holographic boundary Wiggly boundary curve  $(F(t), Z(t) = \epsilon F'(t))$ 



Consider an infalling matter pulse in Poincaré frame  $T_{vv}(t) = E_0 \delta(t)$ , matter (= 2d CFT) quantum effects included Engelsöv-TM-Verlinde '16

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#### Application: Semi-classical evaporation

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 $T_{vv}(t) = 0$  (to allow evaporation)  
 $\Rightarrow E(t) = E_0 e^{-\frac{c}{24\pi C}t} \Rightarrow F(t) = \frac{2}{\alpha A} \frac{h(\alpha)K_0(\alpha e^{-At/2}) - K_0(\alpha)h_0(\alpha e^{-At/2})}{h_1(\alpha)K_0(\alpha e^{-At/2}) + K_1(\alpha)h_0(\alpha e^{-At/2})}$   
where  $A = \frac{c}{24\pi C}, \ \alpha = \frac{24\pi}{c} \sqrt{2CE_0}$ 

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We need to take this seriously Almheiri-Engelhardt-Marolf-Maxfield '19, Penington '19  $\rightarrow$  Jump in minimal quantum extremal (RT) surface (Requires non-perturbative (in  $G_N$ ) knowledge)  $\Rightarrow$  Island rule (see Lárus Thorlacius talk on Friday!) Next: Transfer to Euclidean thermal theory and obtain boundary correlation functions of JT gravity / Schwarzian QM:

$$\langle \mathcal{O}_{h_1} \mathcal{O}_{h_2} \dots \rangle_{\beta} = \frac{1}{Z} \int_{\mathcal{M}} [\mathcal{D}f] \mathcal{O}_{h_1} \mathcal{O}_{h_2} \dots e^{C \int_0^\beta d\tau \{F, \tau\}}$$

with 
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 $f(\tau)$  is reparametrization of  $S^1$ :



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$$Z(\beta) = \left(\frac{2\pi C}{\beta}\right)^{3/2} \exp\left(\frac{2\pi^2 C}{\beta}\right) \sim \int_0^{+\infty} dE \sinh\left(2\pi\sqrt{2CE}\right) e^{-\beta E}$$
  
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Density of states  $\rho(E) = \sinh\left(2\pi\sqrt{2CE}\right)$   
Thermodynamic limit (saddle): found by  $\frac{\partial S}{\partial E} = \beta$  at large  $E$   
 $\implies T \sim \sqrt{E}$ 

 $\rightarrow$  matches semi-classical JT black hole first law, saddle approximation invalid at small *E* 

# Boundary bilocal operator

Take massive scalar field in bulk, asymptotic expansion  $(AdS_2/CFT_1)$ :

$$\phi(Z,F) \quad \to Z^{1-h} \tilde{\phi}_b(F) = \epsilon^{1-h} F'^{1-h} \tilde{\phi}_b(F(\tau)) = \epsilon^{1-h} \phi_b(\tau)$$

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Generating functional:

$$I \sim \int dF_1 \int dF_2 \frac{1}{(F_1 - F_2)^{2h}} \tilde{\phi}_b(F_1) \tilde{\phi}_b(F_2)$$
  
=  $\int d\tau_1 \int d\tau_2 \frac{F'(\tau_1)^h F'(\tau_2)^h}{(F(\tau_1) - F(\tau_2))^{2h}} \phi_b(\tau_1) \phi_b(\tau_2)$ 

Bilocal operator:

$$\mathcal{O}_{h}(\tau_{1},\tau_{2}) \equiv \left(\frac{F'(\tau_{1})F'(\tau_{2})}{(F(\tau_{1})-F(\tau_{2}))^{2}}\right)^{h} \equiv \left(\frac{f'(\tau_{1})f'(\tau_{2})}{\frac{\beta}{\pi}\sin^{2}\frac{\pi}{\beta}[f(\tau_{1})-f(\tau_{2})]}\right)^{h}$$

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$$\mathcal{O}_{h}(\tau_{1},\tau_{2}) \equiv \left(\frac{F'(\tau_{1})F'(\tau_{2})}{(F(\tau_{1})-F(\tau_{2}))^{2}}\right)^{h} \equiv \left(\frac{f'(\tau_{1})f'(\tau_{2})}{\frac{\beta}{\pi}\sin^{2}\frac{\pi}{\beta}[f(\tau_{1})-f(\tau_{2})]}\right)^{h}$$

#### Other origins of this operator:

Boundary-anchored Wilson line in 1<sup>st</sup> order SL(2, R) gauge formulation of JT gravity Blommaert-TM-Verschelde '18, Iliesiu-Pufu-Verlinde-Wang '19

#### Thomas Mertens

Several approaches to obtain JT disk amplitudes exist:

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- ► 1d Liouville Bagrets-Altland-Kamenev '16, '17
- ► 2d Liouville CFT between ZZ-branes TM-Turiaci-Verlinde '17, TM '18
- ▶ 2d BF bulk Blommaert-TM-Verschelde '18, Iliesiu-Pufu-Verlinde-Wang '19 Using first-order formulation of JT gravity in terms of an  $SL(2, \mathbb{R})$  BF gauge theory
- ► Particle in infinite B-field in AdS<sub>2</sub> Yang '18, Kitaev-Suh '18, Suh '20
- Liouville gravity / minimal string TM-Turiaci '19, '20, TM '20

#### Real-time two-point function

Real-time two-point function  $G_h(t) = \langle \mathcal{O}_h(t,0) \rangle$ =  $\frac{1}{Z(\beta)} \int d\mu(E_1) d\mu(E_2) e^{it(E_1 - E_2) - \beta E_2} \frac{\Gamma(h \pm i\sqrt{2CE_1} \pm i\sqrt{2CE_2})}{\Gamma(2h)}$ with  $d\mu(E) = dE \sinh 2\pi\sqrt{2CE}$ 

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#### Behavior full quantum expression:

Always decaying, where late-time behavior of correlator gets strong QG fluctuations (power-law decay  $\sim 1/t^3$  instead of exponential decay)

#### Complexity $\mathcal C$ of boundary theory at time t

= volume of extremal (maximal) surface anchored at the boundary

at points at time t Susskind '14...

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Q: Complexity plateau? Higher topology / random matrices?

Schwarzian out-of-time ordered (OTOC) 4-point function:  $\langle V_1 W_3 V_2 W_4 \rangle = \langle \mathcal{O}_h(t_1, t_2) \mathcal{O}_h(t_3, t_4) \rangle_{OTO}$ 

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Gravitational semi-classical interpretation: OTOC  $\langle V_1 W_3 V_2 W_4 \rangle$  in boundary theory captures gravitational shockwave behavior Shenker-Stanford 15

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Large *C* limit of complete OTOC with light *h*, gives full Dray-'t Hooft eikonal shockwave expressions Lam-TM-Turiaci-Verlinde '18

# Path integral (in Euclidean signature): $\int \frac{[\mathcal{D}g_{\mu\nu}]}{\text{Vol}(\text{Diff})} [\mathcal{D}\Phi] e^{-S_{\text{JT}}-S_{bdy}} = \int \frac{[\mathcal{D}g_{\mu\nu}]}{\text{Vol}(\text{Diff})} \delta(R-\Lambda) e^{-S_{bdy}}$

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Manifold with holographic boundary: non-trivial boundary dynamics, described by Schwarzian action Path integral (in Euclidean signature):

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- Manifold with holographic boundary: non-trivial boundary dynamics, described by Schwarzian action
- Manifold with only geodesic boundaries (K = 0, S<sub>bdy</sub> = 0): remaining path integral ∫<sub>R=Λ</sub> 1/√ol(Diff) [Dg<sub>µν</sub>] × 1 is computing volume integral of all inequivalent metrics one can put on a 2d manifold with given R = Λ and geodesic boundaries → Mathematicians call this the volume of moduli space of Riemann surfaces = Weil-Petersson volume

In order to analyze higher topology, we augment the action by the Einstein-Hilbert action:

$$\begin{split} S &= -\frac{S_0}{2\pi} \left[ \frac{1}{2} \int R + \oint K \right] + S_{JT} \\ &= -\chi S_0 + S_{JT} \text{ with } \chi \text{ the Euler characteristic: } \chi = 2 - 2g - n \text{ for} \end{split}$$

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#### Motivation:

Comes from zero'th order term in near-extremal near-horizon expansion of higher dim. black holes where  $S_0$  is the extremal entropy

## Multiboundary amplitudes (3)

Multiboundary amplitudes Saad-Shenker-Stanford '19:



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Ingredients:

- Single-trumpet  $Z_{\rm JT}(\beta, b) \sim \beta^{-1/2} e^{-\frac{1}{4\beta}b^2}$
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Weil-Petersson (WP) volume V<sub>g,n</sub>(b) ≡ V<sub>g,n</sub>(b<sub>1</sub>...b<sub>n</sub>) volume of moduli space of Riemann surfaces of genus g with n geodesic boundaries of length b<sub>i</sub> multivariate polynomials in b<sup>2</sup><sub>i</sub>

## JT gravity as a matrix integral

#### Mirzakhani '07:

Weil-Petersson volumes satisfy recursion relations

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#### Eynard-Orantin '07:

→ match with topological recursion relations of a double-scaled matrix model with leading (large *L*) spectral density  $\langle \rho(E) \rangle_0 = \rho_0(E) = L \sinh 2\pi \sqrt{E}$ 

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Multi-loop correlator  $\langle Z(\beta_1) \dots Z(\beta_n) \rangle$  in matrix integral with above  $\rho_0(E)$ , identifiable as multi-boundary amplitude in gravity,  $\langle \dots \rangle = \int dH \dots \exp(-\operatorname{Tr} V(H))$ 

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Double-scaled expansion parameter L matches with  $e^{S_0}$ Holographic interpretation:

- $\rightarrow$  Boundary Hamiltonian  $\equiv$  random matrix H
- $\rightarrow$  Ensemble-averaged holography (cfr. SYK has some averaging)

Boundary two-point function in thermal AdS/CFT decays exponentially at late times due to bulk quasinormal modes: E.g. in 1+1d:  $\langle \mathcal{O}(t)\mathcal{O}(0)\rangle_{\beta} = \frac{1}{\left(\sinh\frac{\pi}{\beta}t\right)^{2h}} \sim e^{-\frac{2\pi}{\beta}ht}$ 

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Spectral form factor: Cotler et al. '16, Saad-Shenker-Stanford '18  $|Z(\beta + it)|^2 = \sum_{n,m} e^{-\beta(E_n + E_m)} e^{it(E_m - E_n)}$ Late-time low mean  $\sim Z(2\beta) \neq 0$ 

Typical form of spectral form factor in (averaged) chaotic systems



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 $\mathsf{JT}$  gravity has such late-time behavior for its spectral form factor

- $\rightarrow$  easily proved using matrix description
- $\rightarrow$  has gravitational interpretation in terms of higher topology

In matrix integral, the spectral form factor is  $\langle Z(\beta + it)Z(\beta - it) \rangle$  (two analytically continued macroscopic loop operators inserted in matrix integral)

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Interpretation: disconnected piece, sine kernel, and contact term

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Universal for (hermitian) random matrix systems away from the spectral edge, and when  $|E - E'| \ll 1$  $\rightarrow$  will hold for JT gravity as well

Factorized term dominates as  $\sim e^{2S_0}$  ( $\rho_0(E) \sim e^{S_0} \sinh 2\pi \sqrt{E}$ ) However, at very late times  $t \sim e^{S_0}$ , one has  $E - E' \sim e^{-S_0}$  and this can compensate the suppression and give important effects!

Now we can understand the late-time behavior of the spectral form factor in JT gravity  ${\tt Saad-Shenker-Stanford}$  '19

Slope: Factorized contribution  $\rho_0(E)\rho_0(E')$ :

$$Z(\beta+it) = e^{S_0} \left(\frac{\pi}{\beta+it}\right)^{3/2} e^{\pi^2/(\beta+it)} \rightarrow |Z(\beta+it)|^2 \sim \frac{e^{2S_0}}{t^3}$$

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 $Z(\beta + it) = e^{S_0} \left(\frac{\pi}{\beta + it}\right)^{3/2} e^{\pi^2/(\beta + it)} \rightarrow |Z(\beta + it)|^2 \sim \frac{e^{2S_0}}{t^3}$ Geometry: Matches with disconnected disk geometries  $\int_{\beta_0}^{\beta_2} \text{Late-time decay of Schwarzian regime}$ 

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Full non-perturbative answer for amplitudes using numerical matrix model techniques Johnson '19-'20

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Reinstating  $\langle n | O | m \rangle$  = study boundary two-point function Leads to similar picture Blommaert-TM-Verschelde '19, Saad '19

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► textbook holography has fixed Hamiltonian: *N* = 4 SYM dual to *AdS*<sub>5</sub> × *S*<sup>5</sup> type IIB → fixed Hamiltonian (no ensemble averaging) → no connected topologies ?

## Jackiw-Teitelboim gravity is toy model of quantum gravity, which is both relevant and solvable

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 $\mathsf{JT}$  gravity is ideal test case to study conceptual questions about quantum gravity